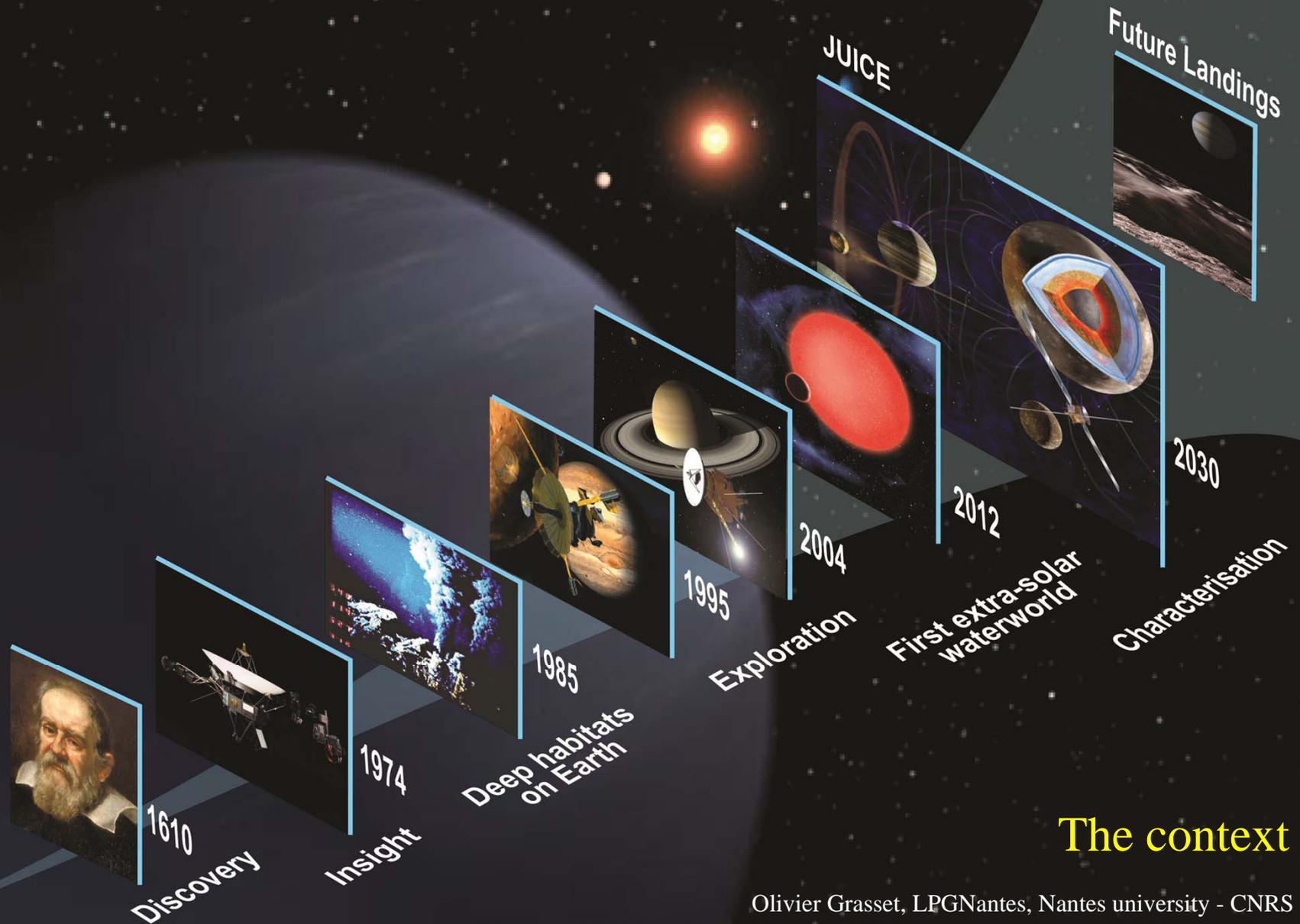


Habitabilité des lunes de Jupiter: des premières évidences aux futures explorations

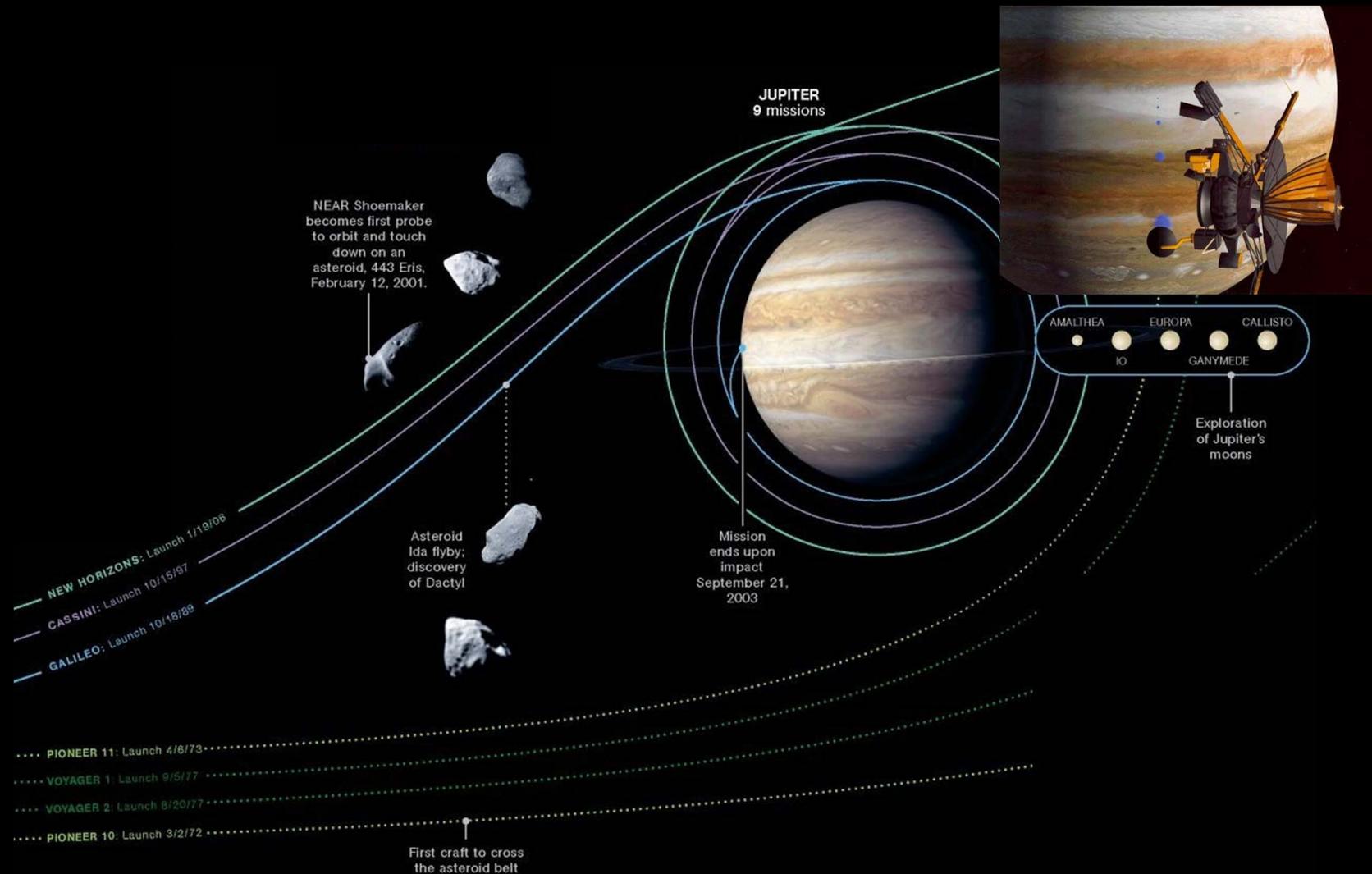


The context

Previous missions

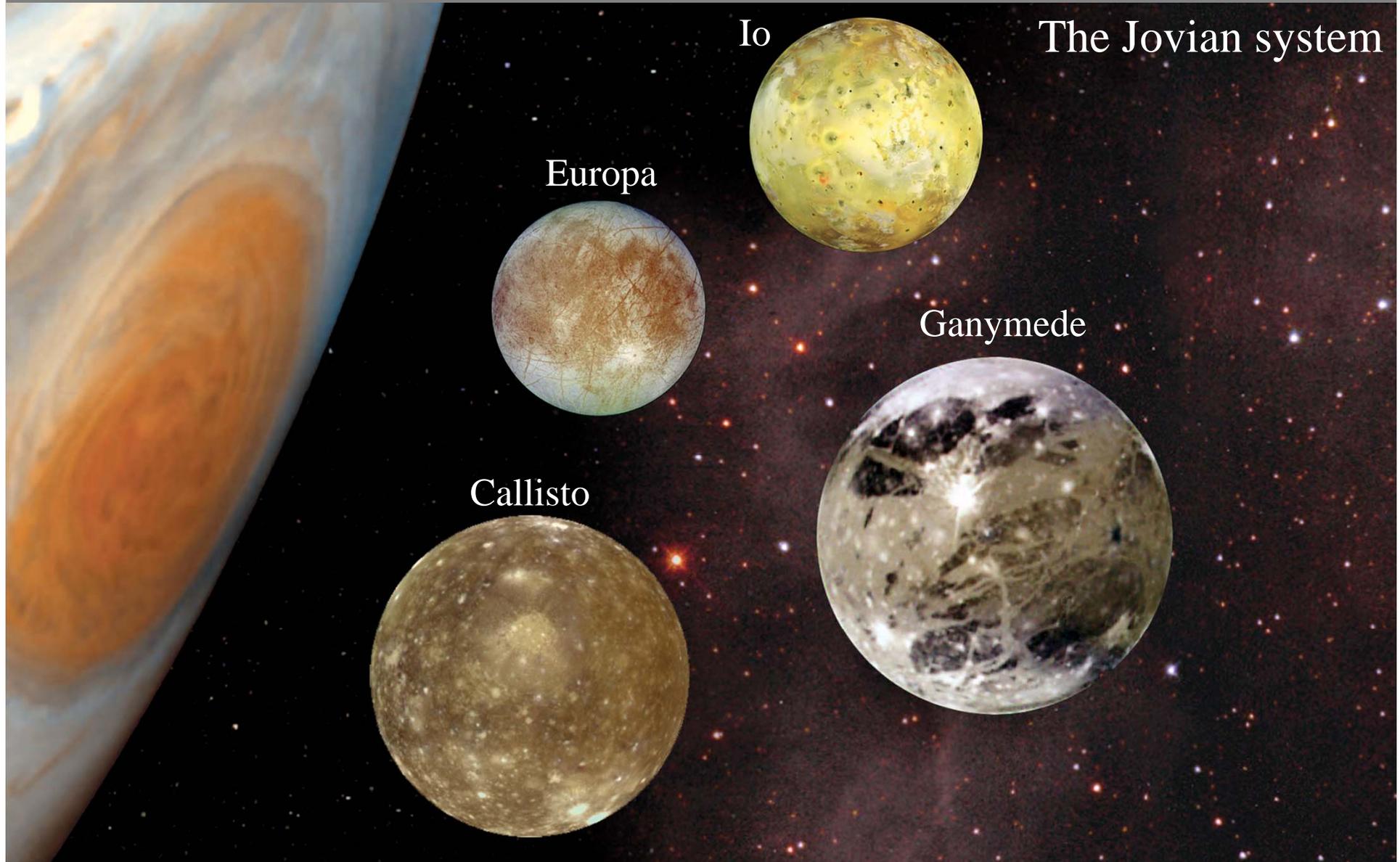
Context

A few flybys and an orbiter (Galileo)



Habitabilité des lunes de Jupiter: des premières évidences aux futures explorations

The planetary bodies that we will explore...

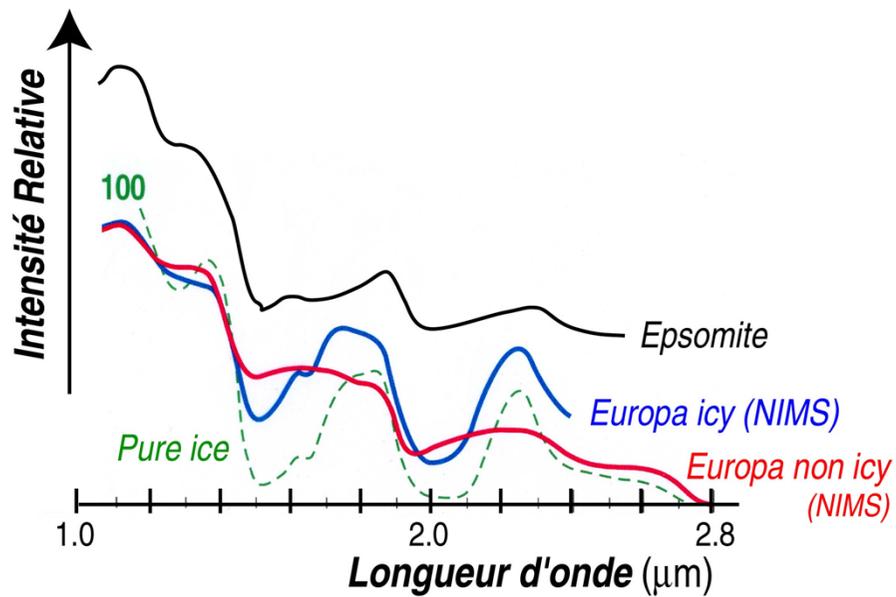


The context

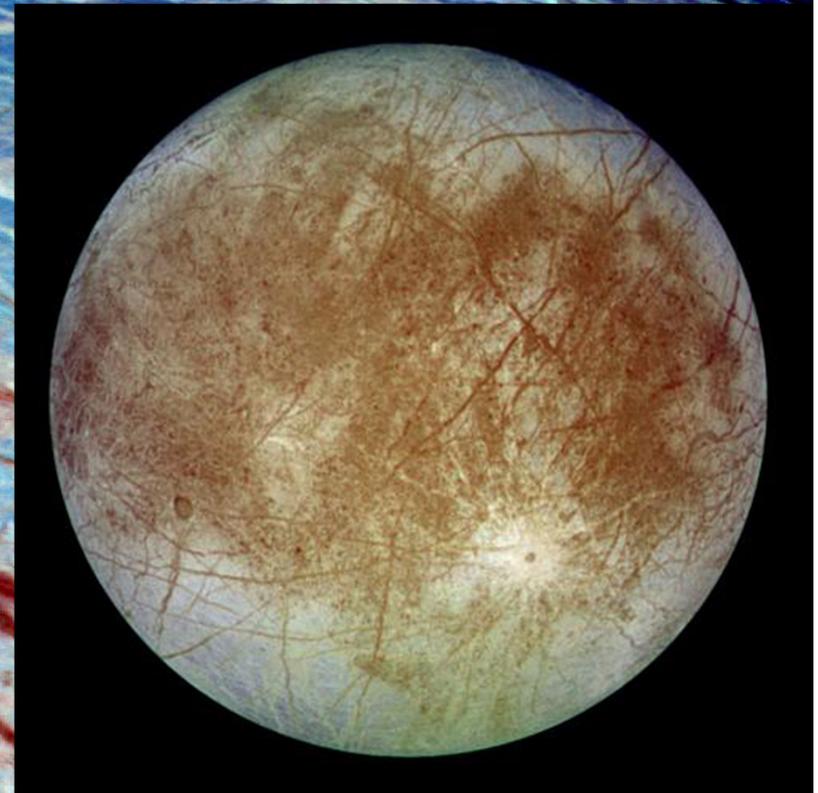
Evidences of possible habitability

Chemical evidences

The composition of ices



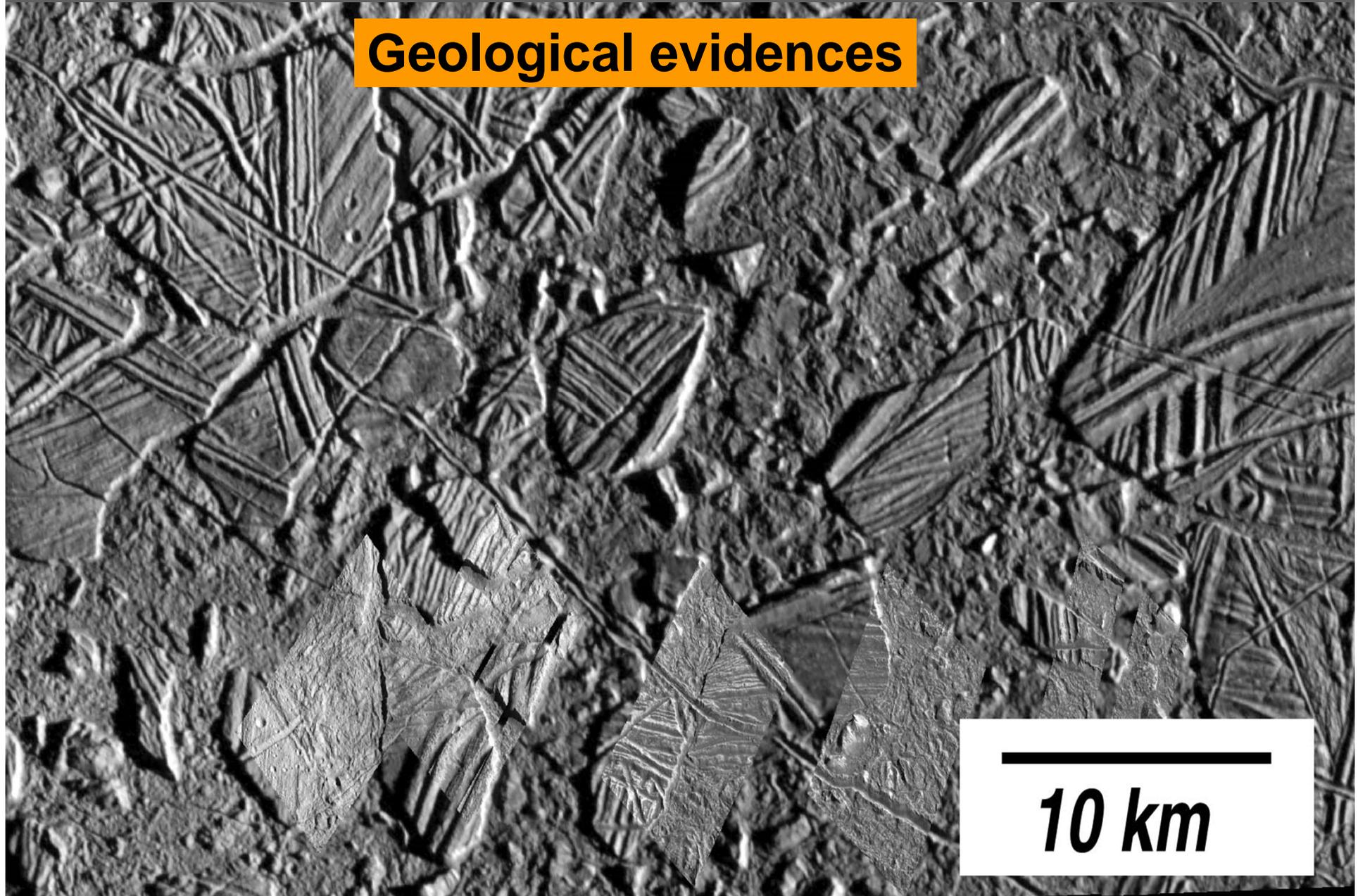
from McCord et al. (1999)



The context

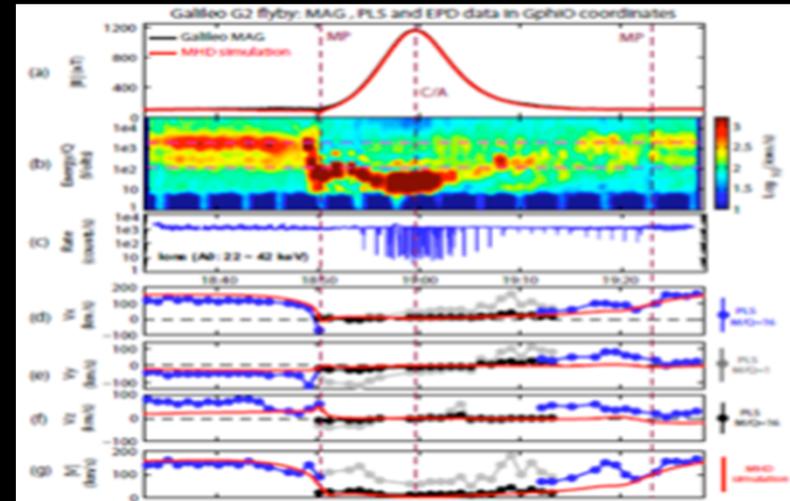
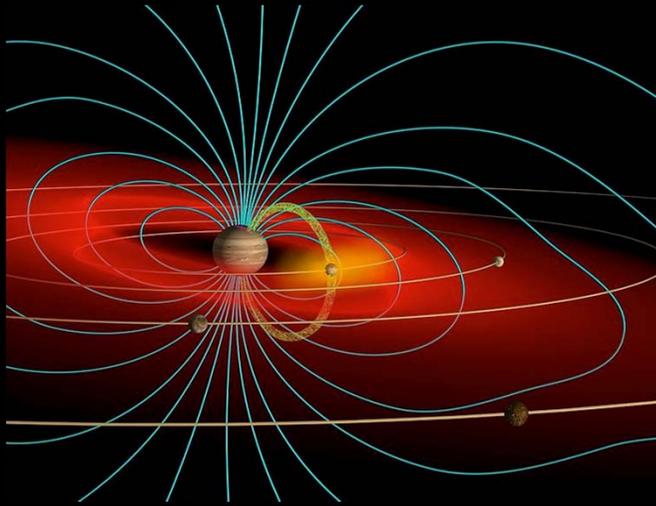
Evidences of possible habitability

Geological evidences

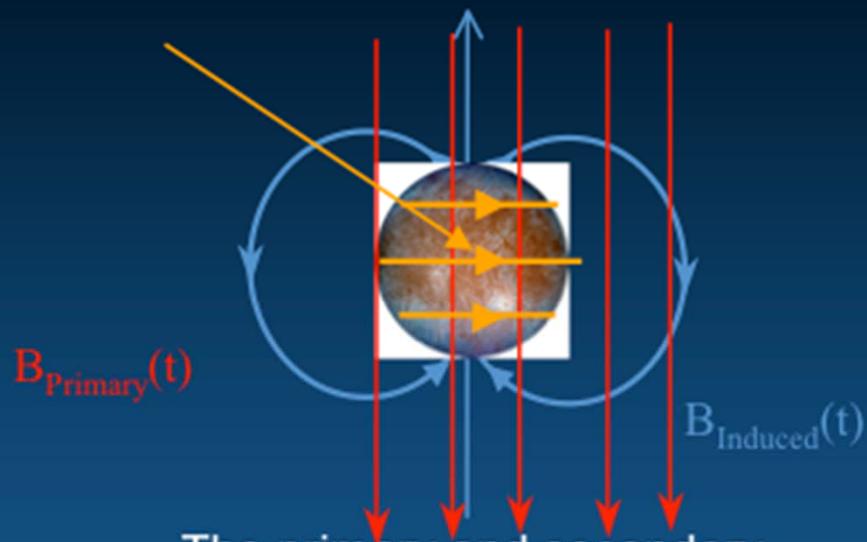


The context

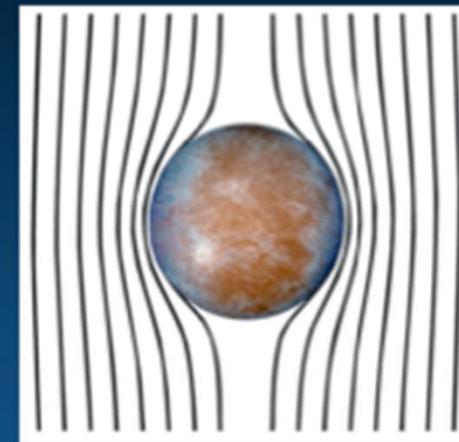
Icy worlds – detection of an induced magnetic field



Eddy currents



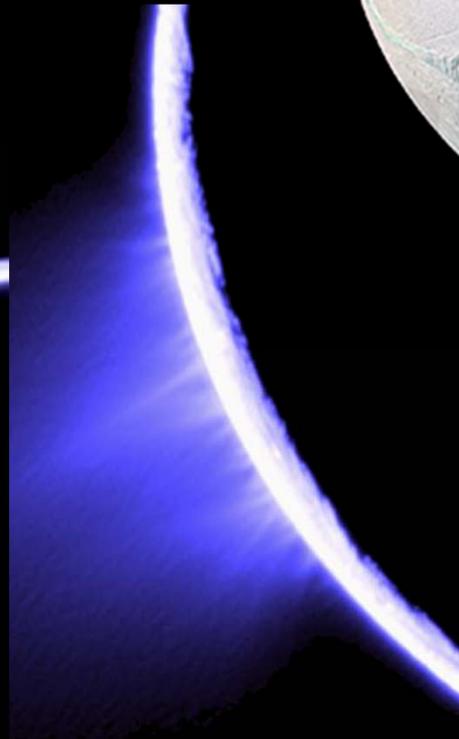
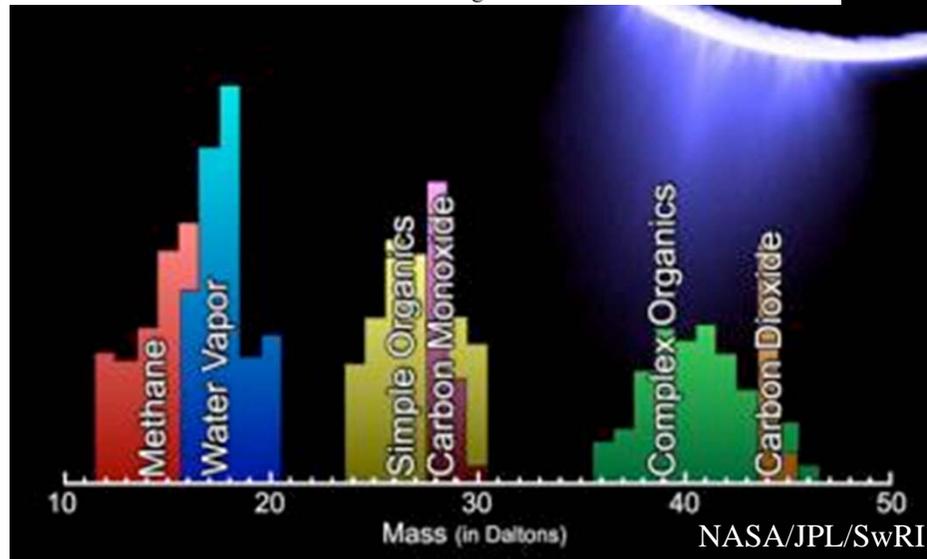
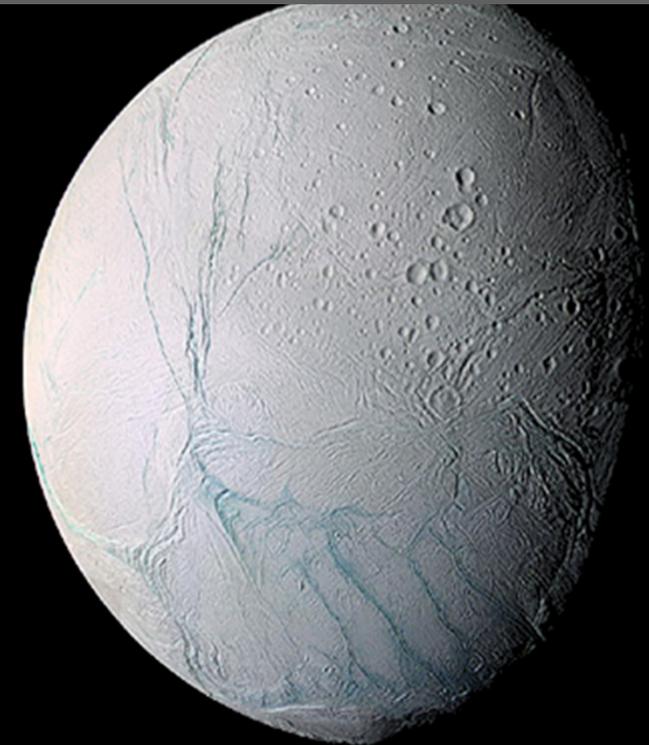
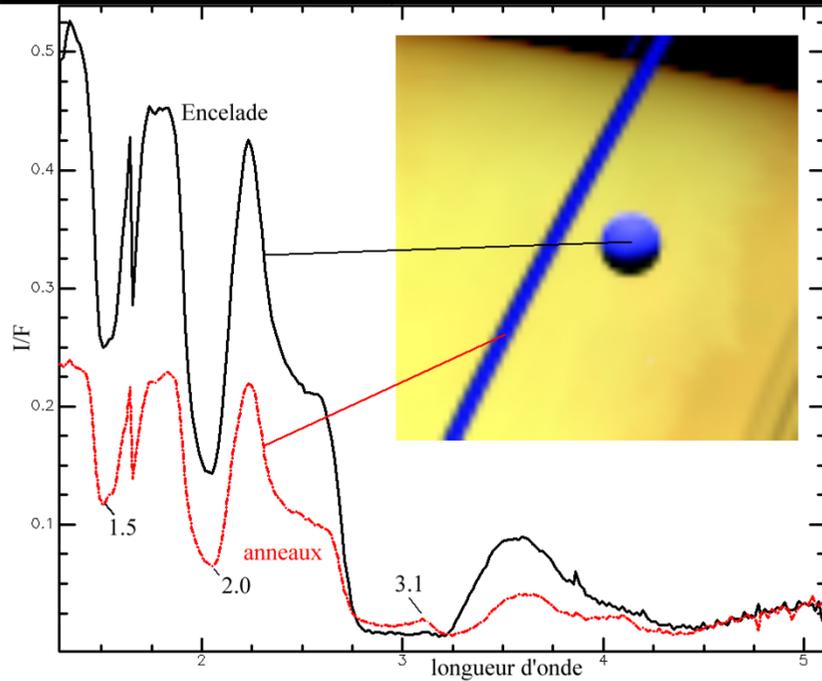
The primary and secondary fields shown separately



The total field

The context

Evidences of possible habitability – Cassini evidences at Enceladus



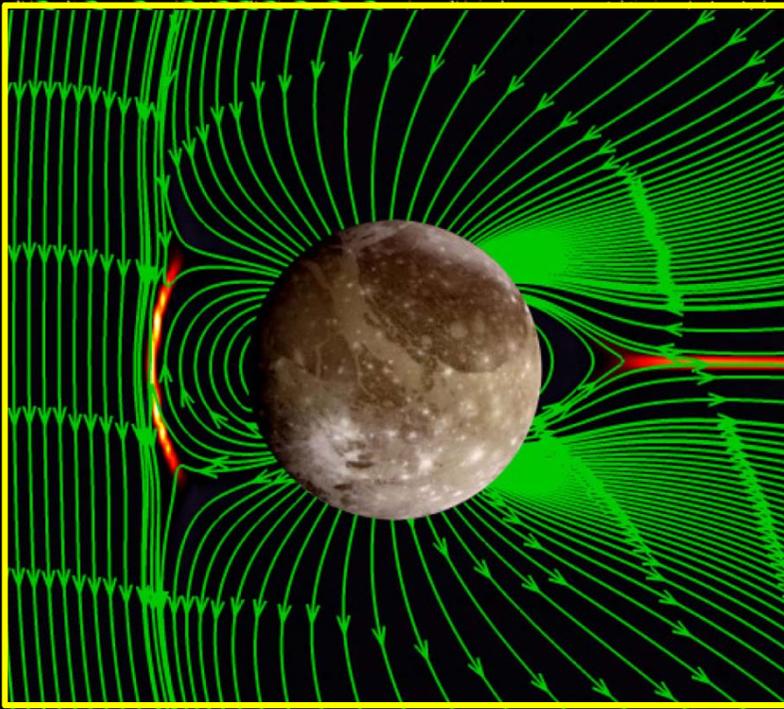
The context

Evidences of possible habitability – Ganymede and Callisto

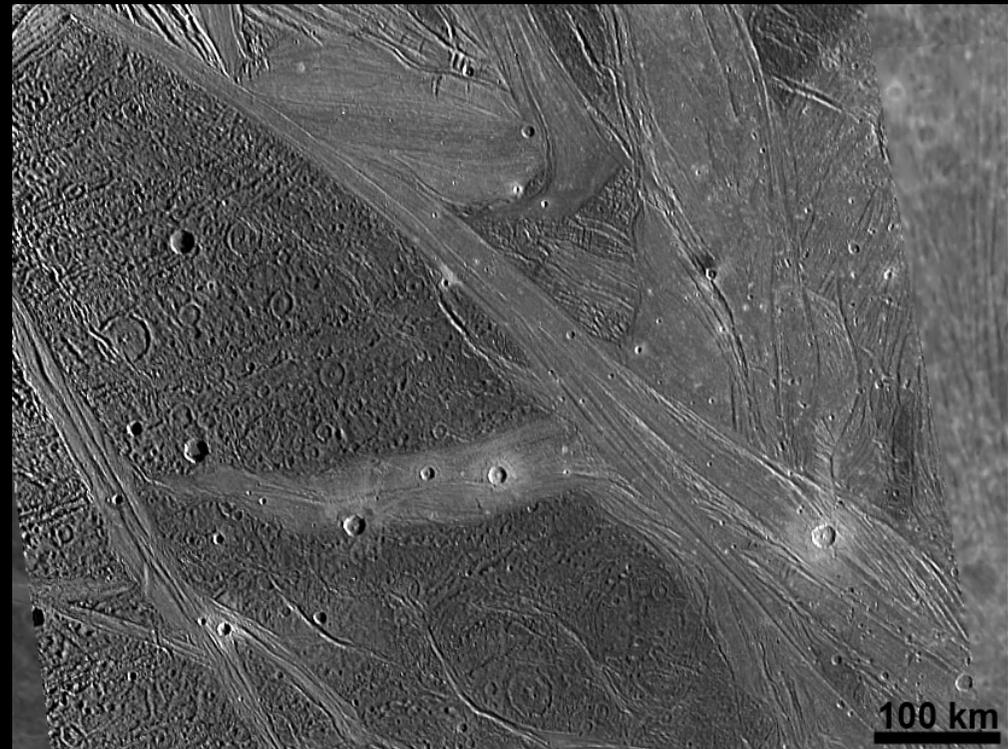
Galileo evidences

Induced magnetic field

Observed but not characterised



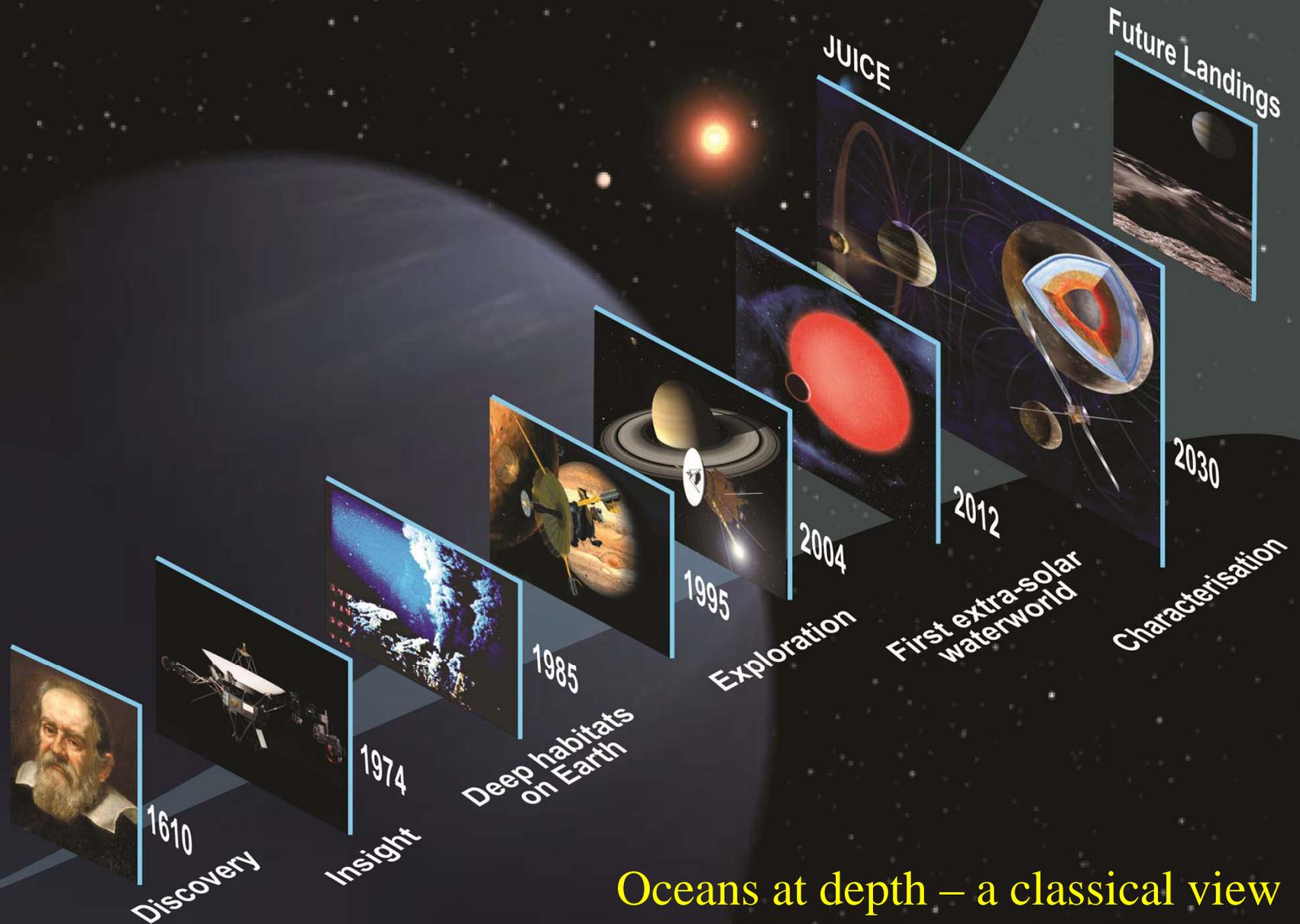
Geologic activity



Questions

- ✧ Which depth?
- ✧ Which size?
- ✧ What is its composition?

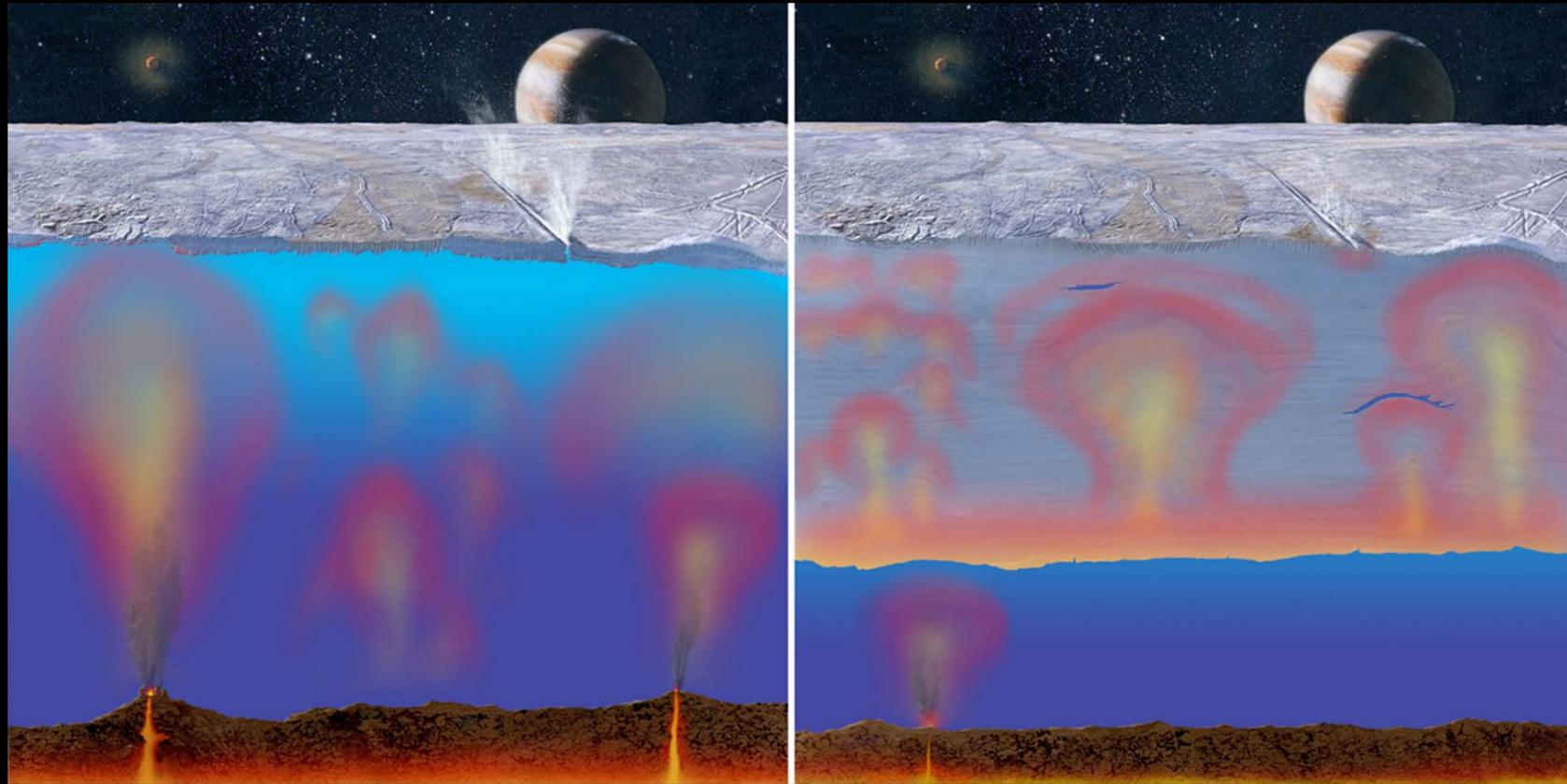
Habitabilité des lunes de Jupiter: des premières évidences aux futures explorations



Oceans at depth – a classical view

Oceans at shallow depth

How deep are the oceans? We still don't know...



Europa

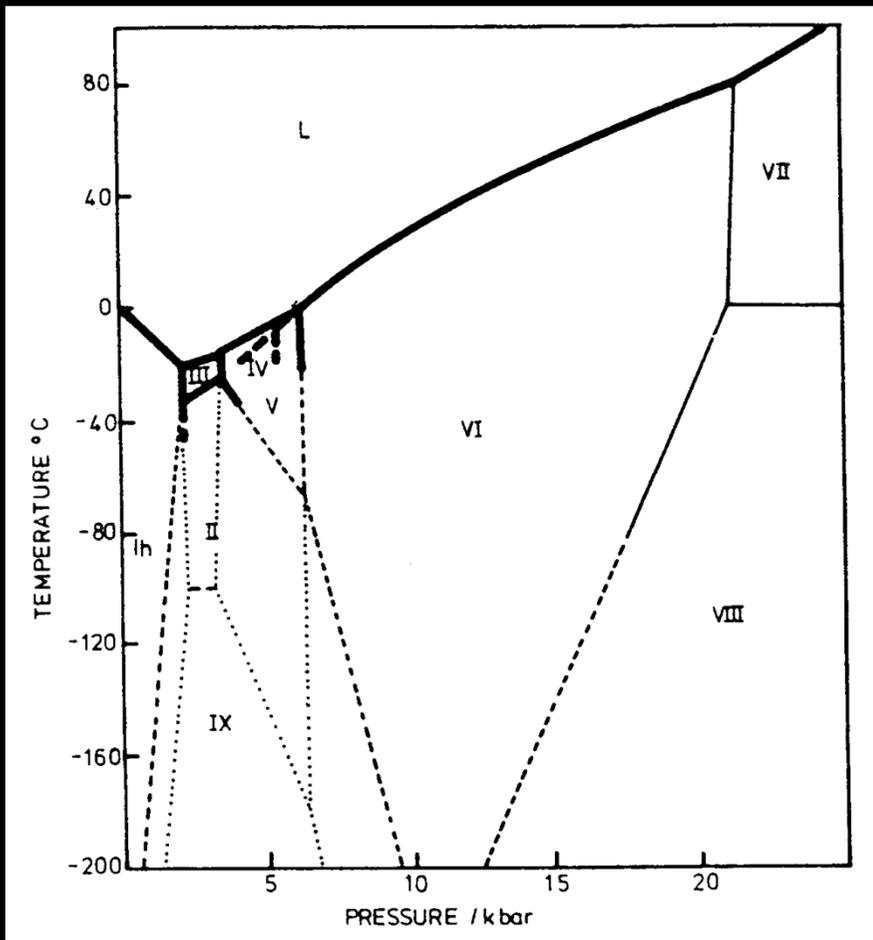
Evidences are not sufficient to solve this issue

Globally thin crust: pros – magnetic field?, geologic features, current activity
cons – thermal equilibrium?, geologic features

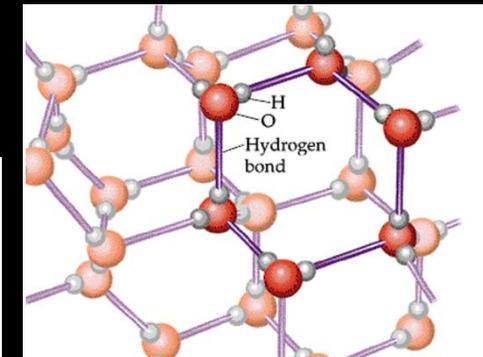
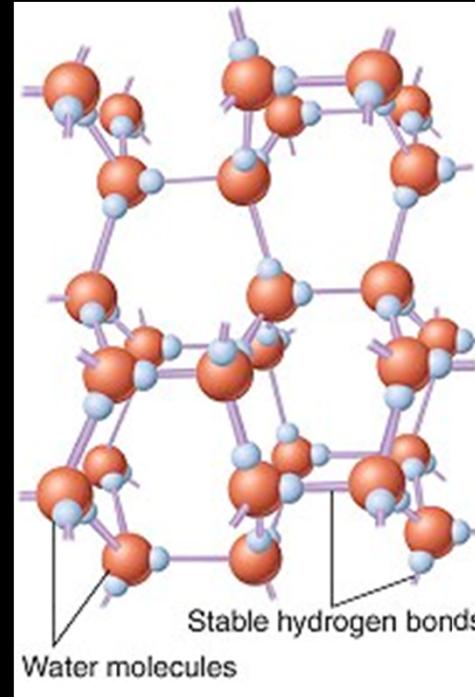
A new space mission is needed

Very deep oceans

Giant Icy worlds – the high pressure phase diagram of water

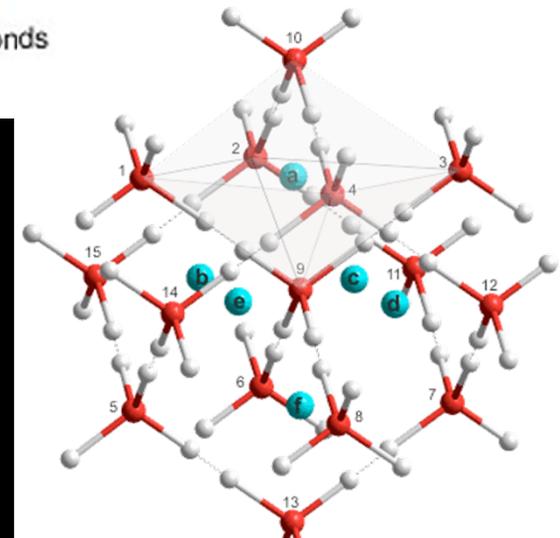


Ice V



Ice Ih

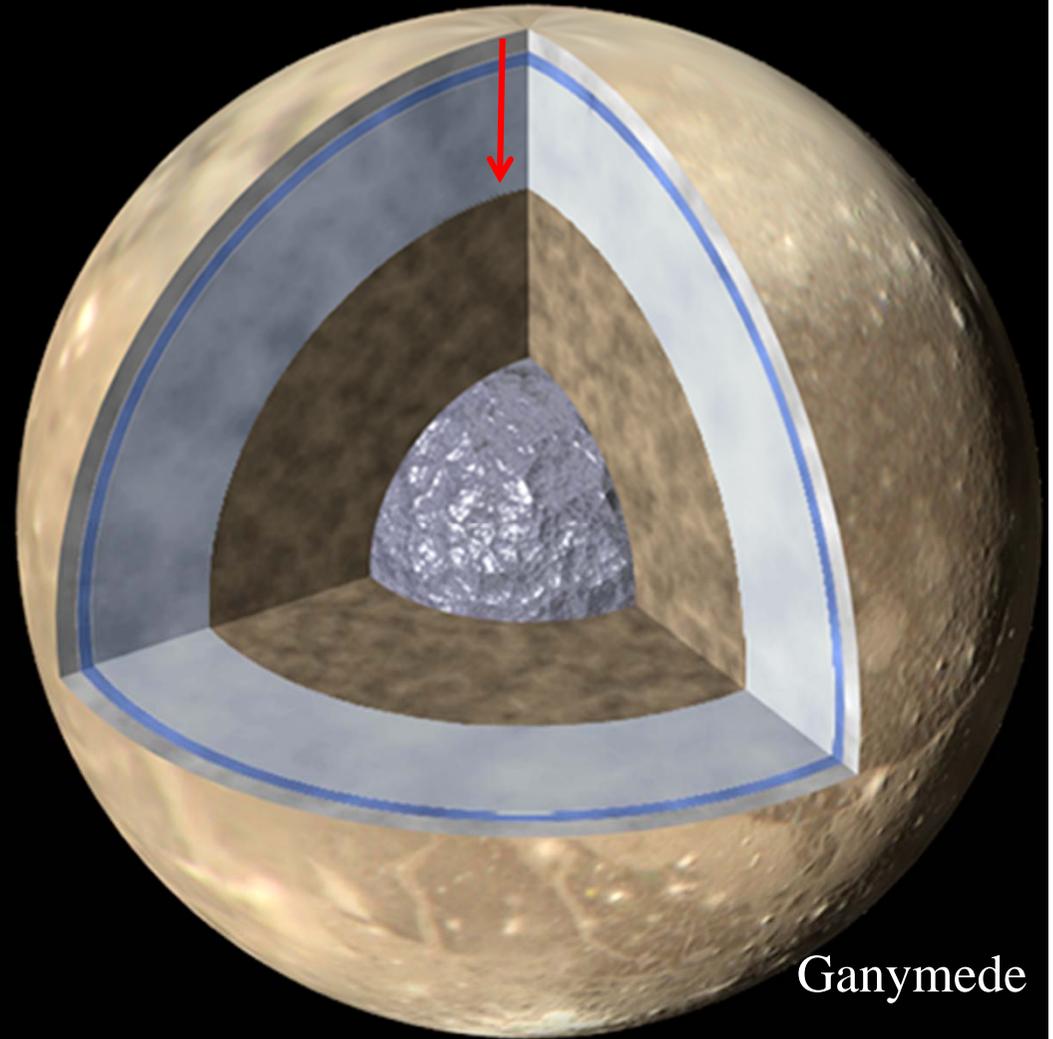
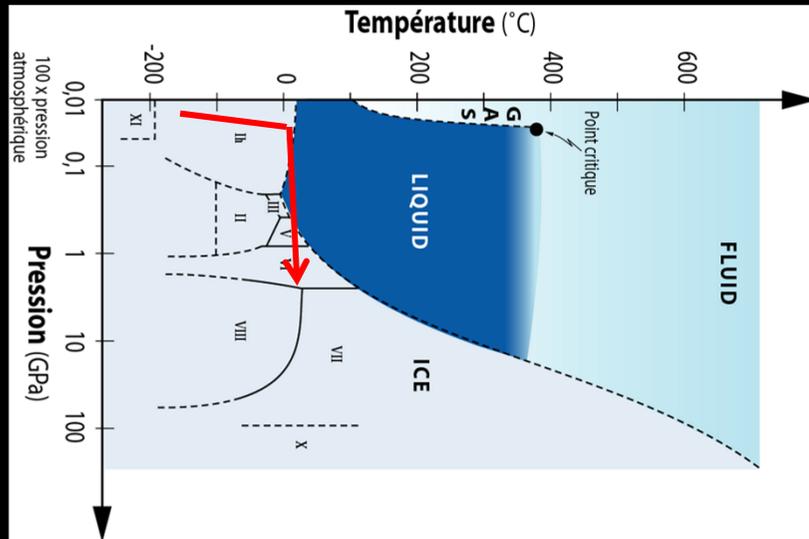
Ice VII



H₂O Well-known since 1912 (Bridgman)
Modern experiments (for planetology) devoted to complex mixtures.

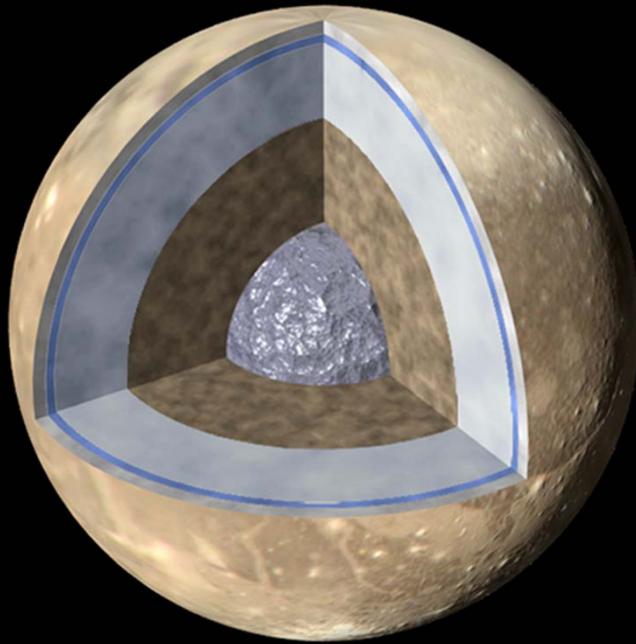
Very deep oceans

Icy worlds – evidences of a liquid ocean

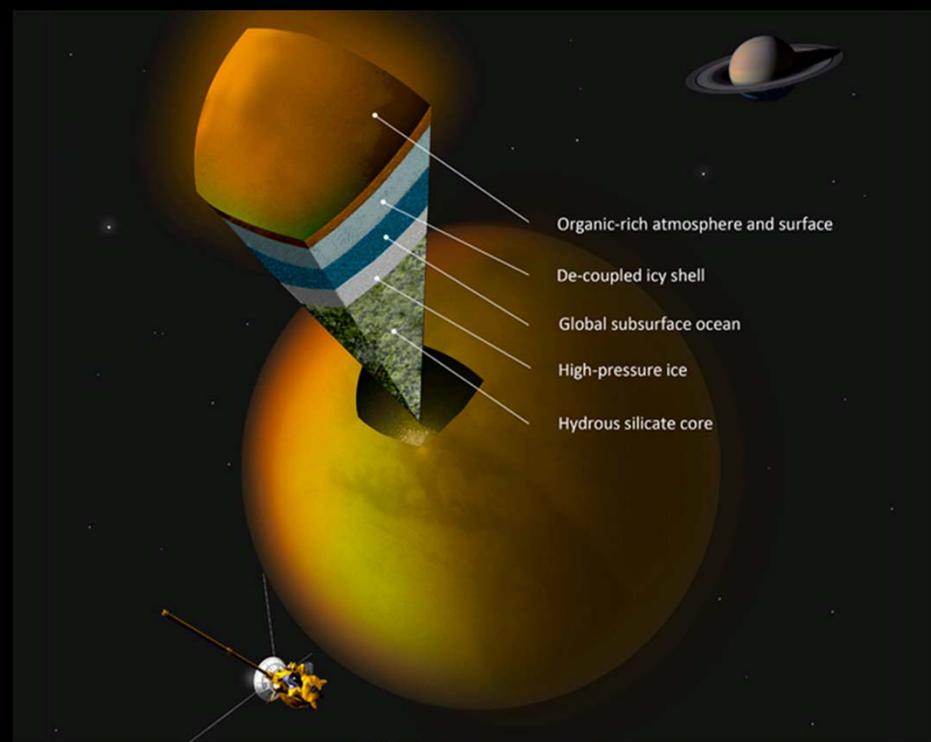


Very deep oceans

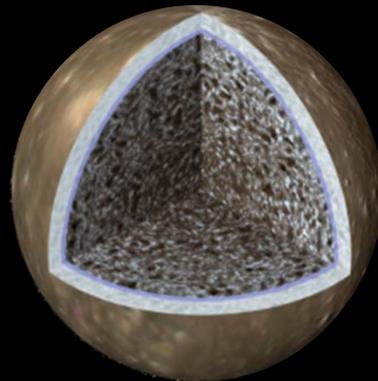
The known examples in the Solar System



Ganymede



Titan



Callisto

Habitabilité des lunes de Jupiter: des premières évidences aux futures explorations

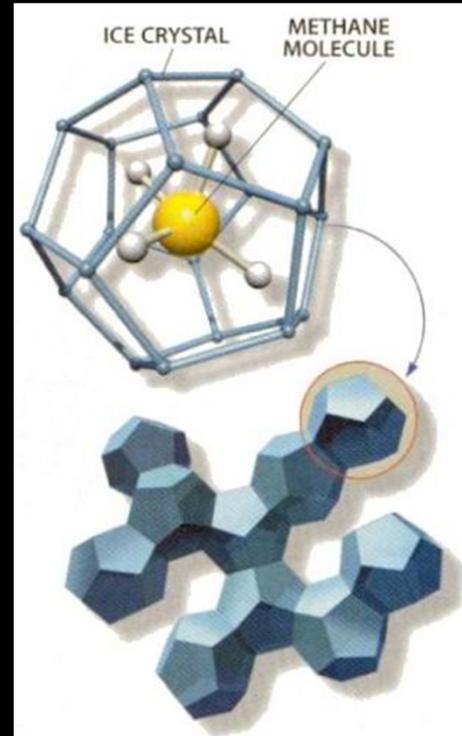
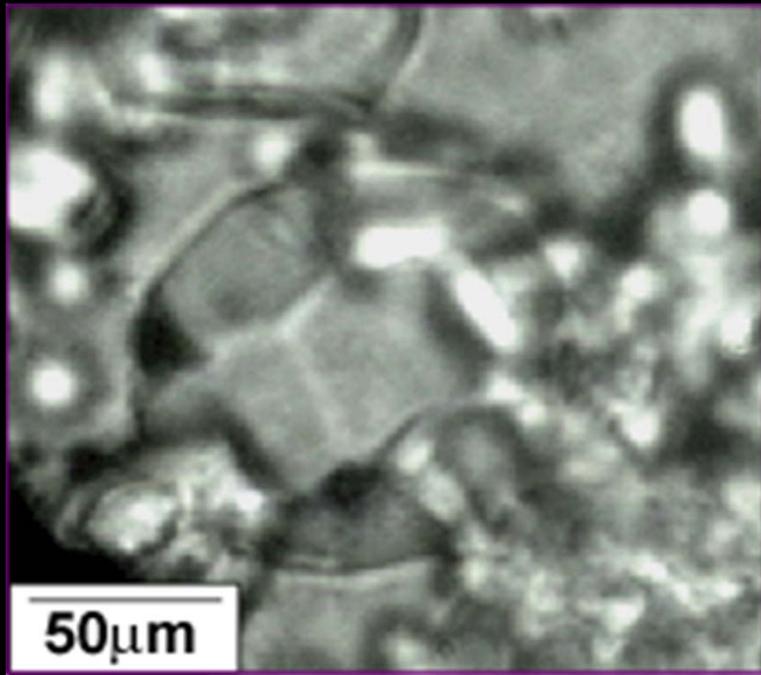


Une vue moins classique

Oceans at shallow depth

Le problème de la composition

Clathrate hydrates : the burning ice.



Known since the 50's (oil engineers)

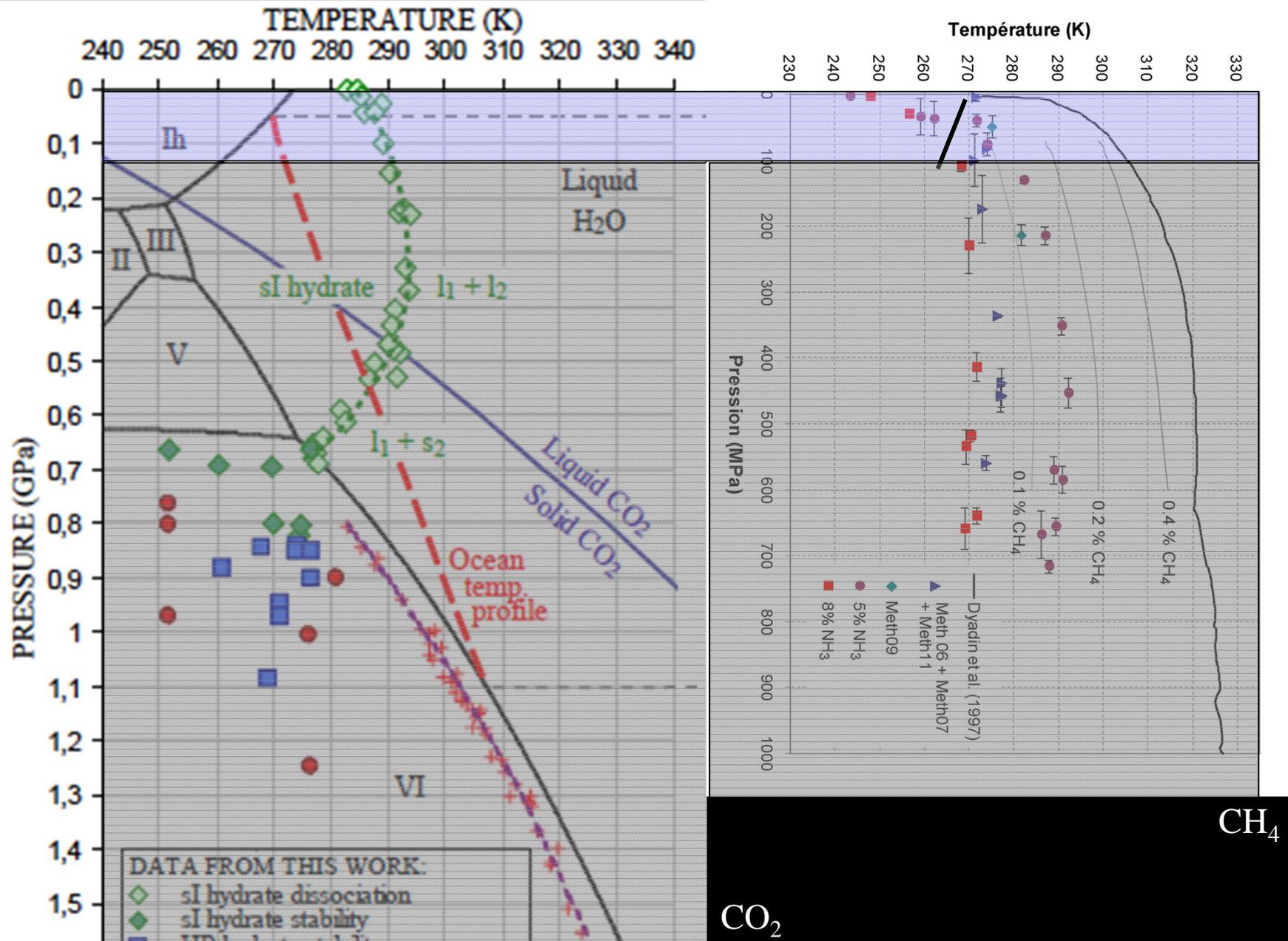
Envisaged in icy moons since early 90s (not sure about this...)

More recently, Prieto-Ballesteros (2005), and Vance et al. envisage to find them in abundance in Europa's crust

But how does that work in depth?

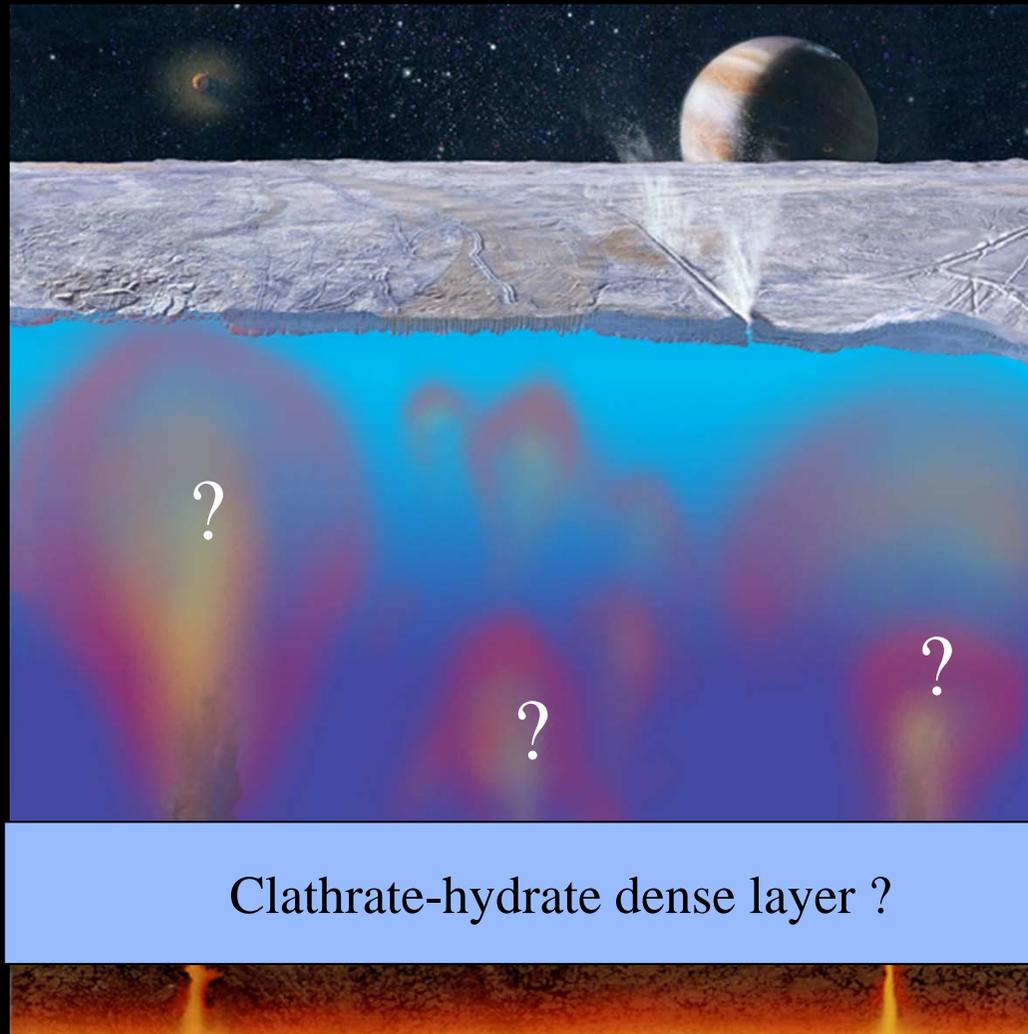
Oceans at shallow depth

Contact of the ocean with silicates - is that certain?



Oceans at shallow depth

Contact of the ocean with silicates - is that certain?



Clathrate-hydrate dense layer ?

At last, the total amount of volatiles and salts fixes the thickness of the deep layer.

But why is that so important? A problem for **SPONCH**.

Oceans at shallow depths

no Phosphorus in planetary ices

Composition moléculaire

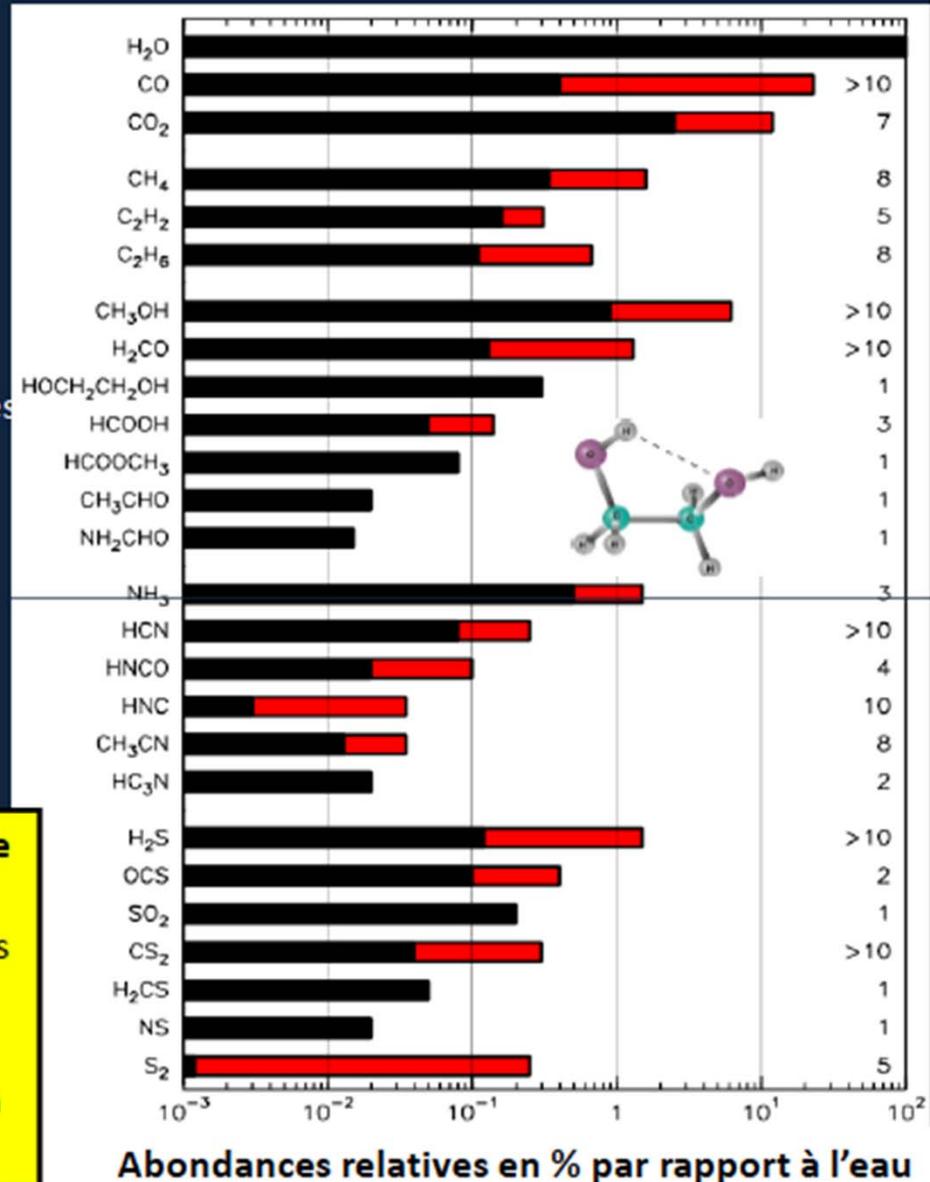
- grande richesse
- molécules complexes (ex: glycol)
- Molécules hydrogénées/oxygénées
- Molécules saturées/non saturées
- 5 ordres de grandeur dans les abondances

Fortes similitudes avec les régions de formation d'étoiles

Formation par les mêmes mécanismes que les molécules interstellaires :

- réactions ion-molécule, et à la surface des grains à basse température

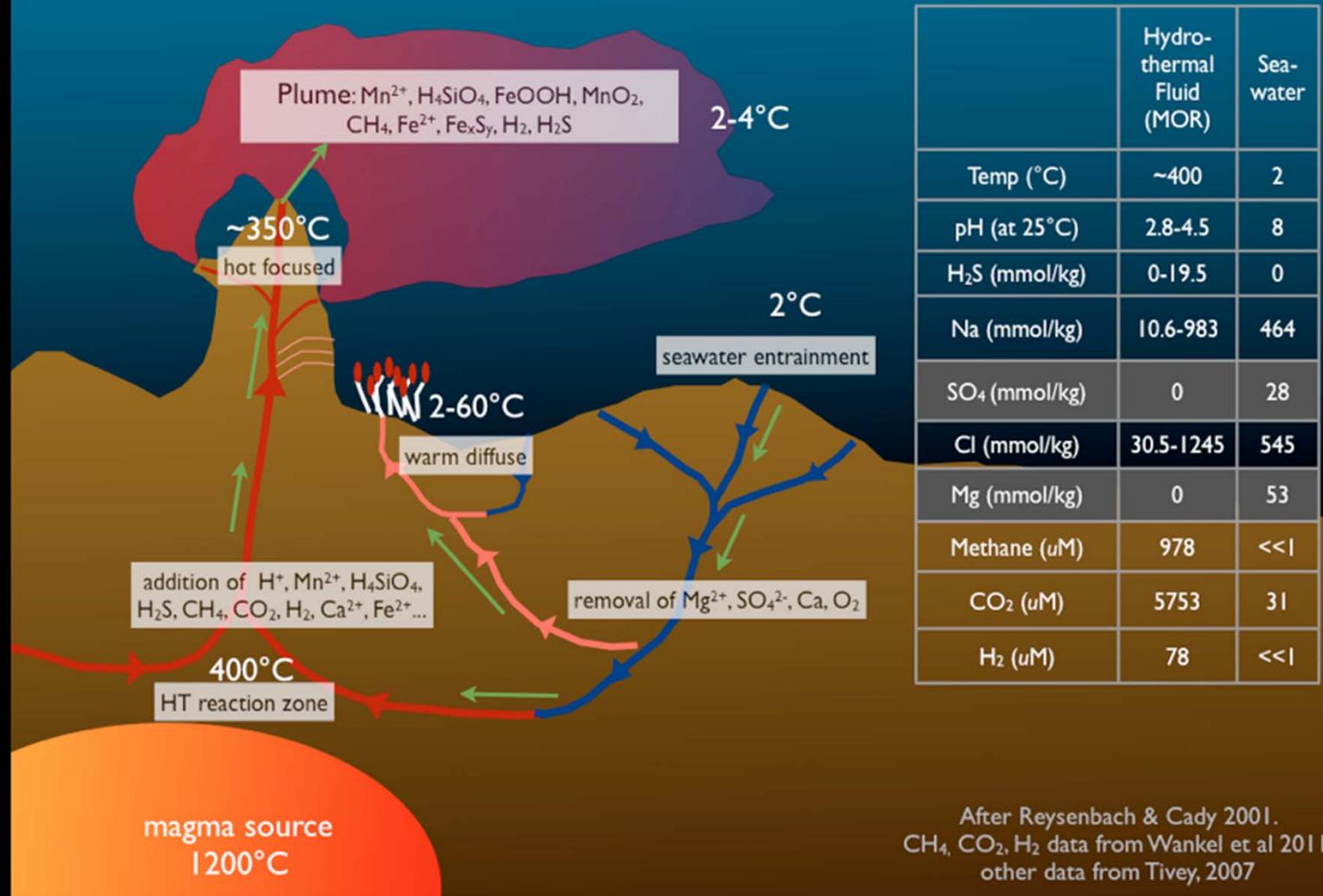
Origine : grains interstellaires ou formation dans les régions extérieures du disque



Oceans at shallow depths

Hydrothermal vents – that helps a lot...

Seawater changes into hydrothermal fluid



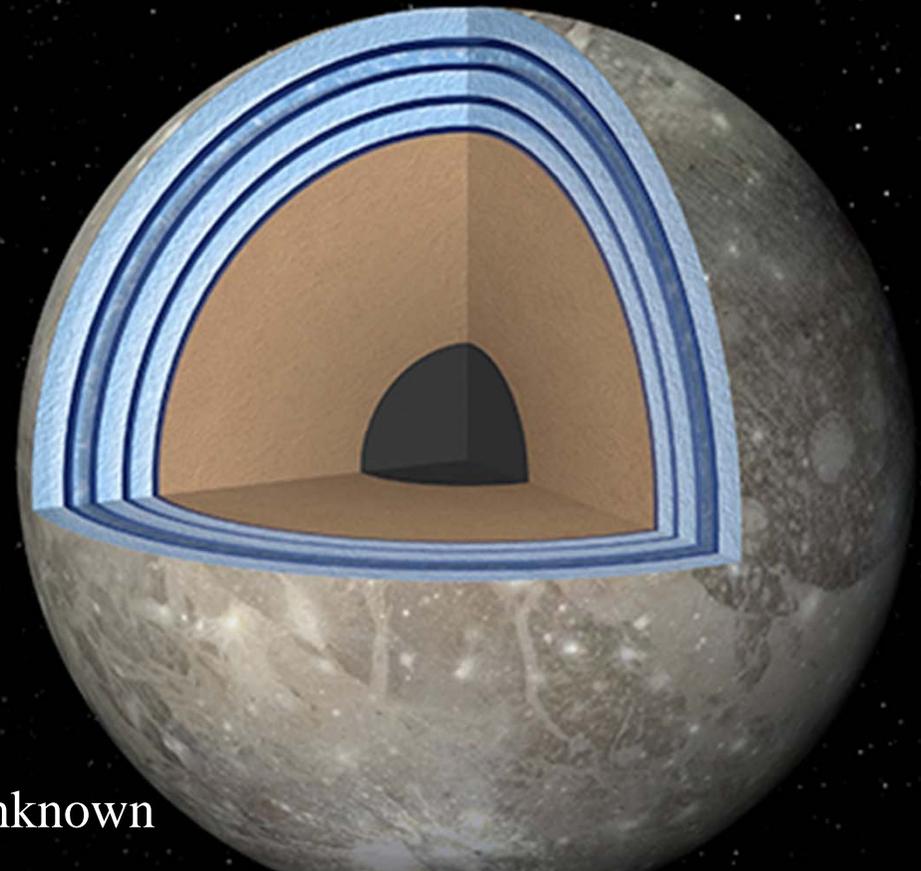
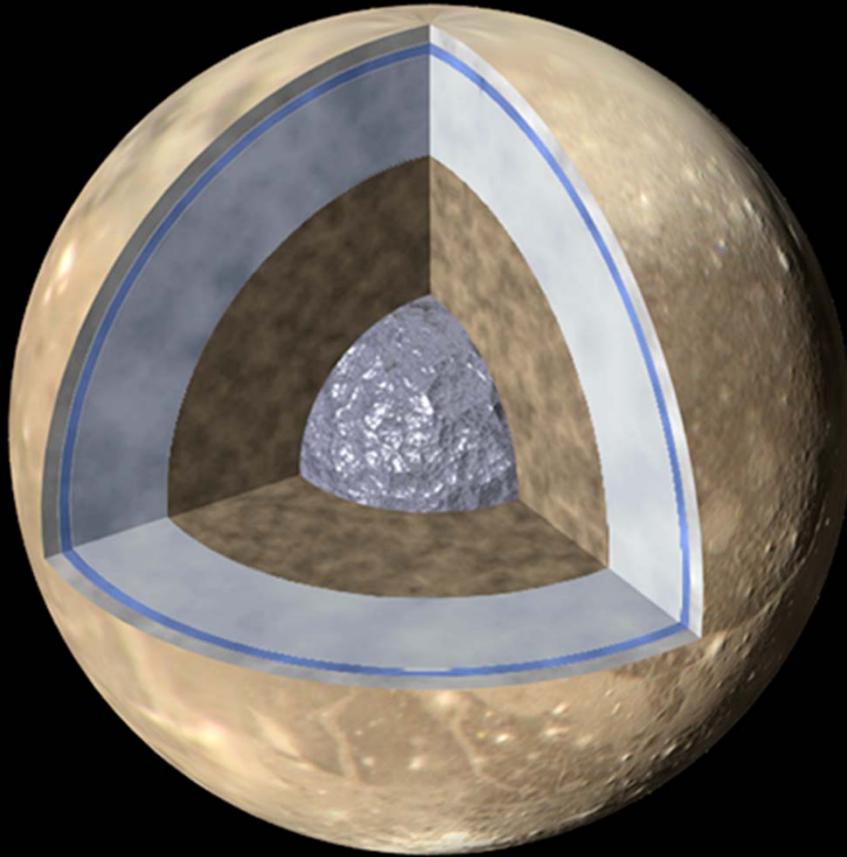
After Reysenbach & Cady 2001.
CH₄, CO₂, H₂ data from Wankel et al 2011,
other data from Tivey, 2007

Very deep oceans

Icy worlds – the complex chemistry is a key player (NH_3 , H_2SO_4 , CO_2 , H_2S , ...)

Density contrasts of water mixtures...

Vance et al., 2014

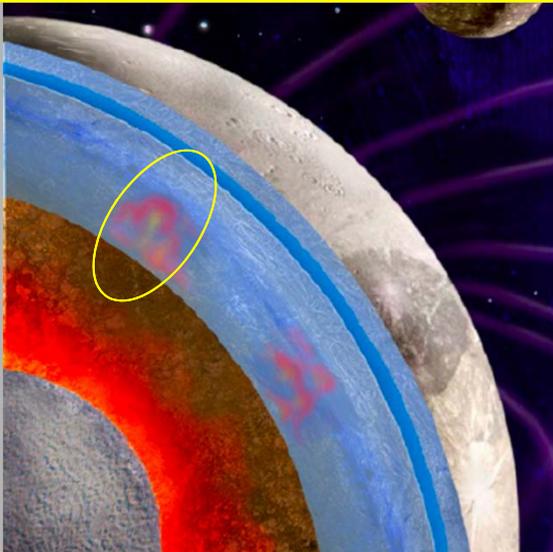


The internal structure of the moons is still unknown

Very deep oceans

A comparison

Ganymede type: liquid layers trapped between two icy mantles



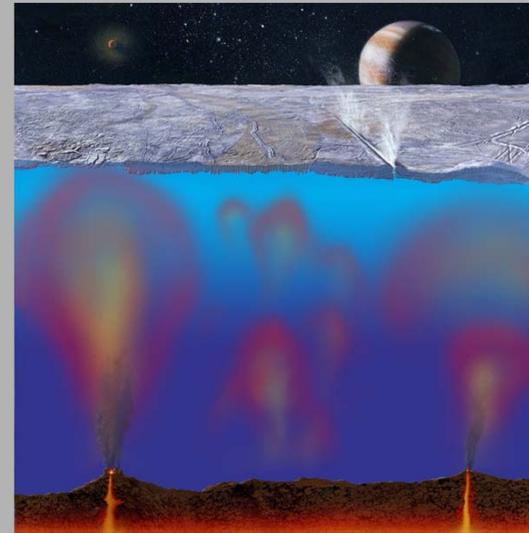
Occurrence:

Largest moons, hot ice giants, ocean-planets...
Most common habitat in the universe ?

Key question:

Are these waterworlds habitable ?

Europa type: Liquid layer in direct contact with silicates (Earth's analog)



Occurrence:

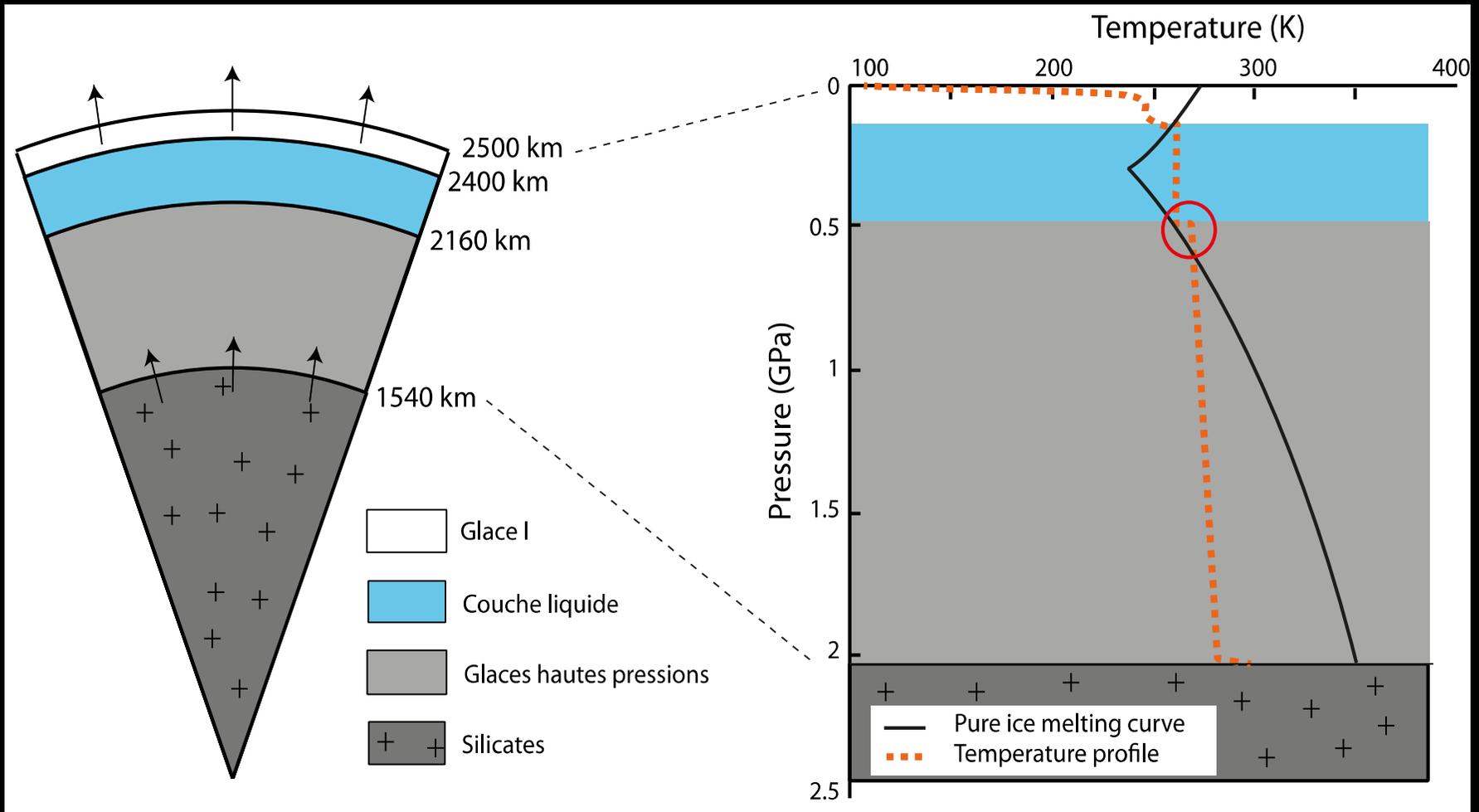
Europa, Enceladus
Only possible for very small bodies

Key questions:

How are the surface active areas related to potential deep habitats?
Is the water in contact with silicates?

Very deep oceans

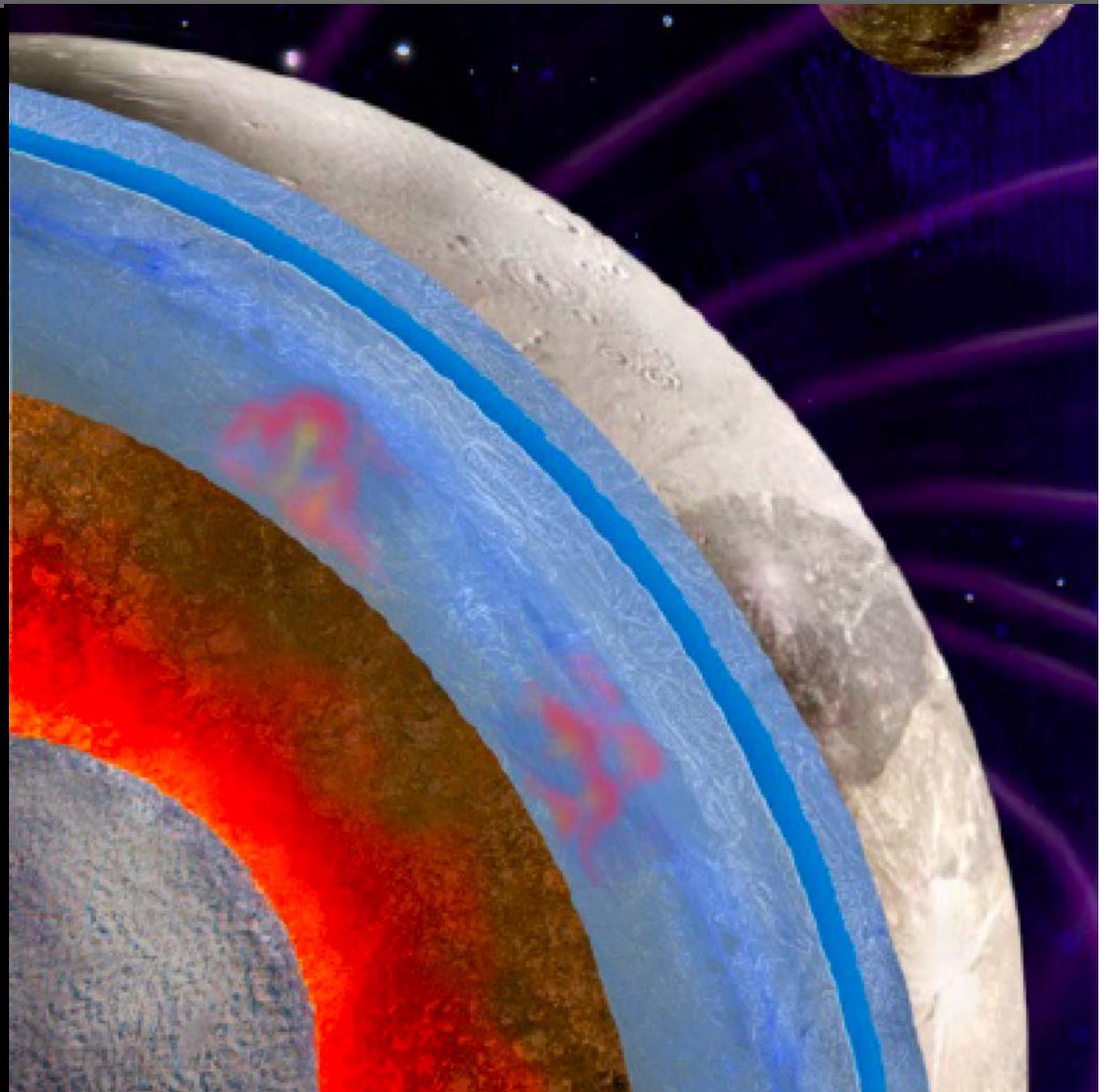
The dynamics of the system



Convective regime

How does that work?
Still unsolved

Very deep oceans



Habitabilité des lunes de Jupiter: des premières évidences aux futures explorations



Future space missions

JUICE: JUpiter Icy moons Explorer

Emergence of habitable worlds around gas giants

Jupiter system as an archetype for gas giants

Callisto:

remnant of the early solar system

- Icy shell, ocean
- Geology, surface composition
- Past activity

Europa: recently active zones

- Surface non-water-ice material
- Search for liquid water
- Recent activity

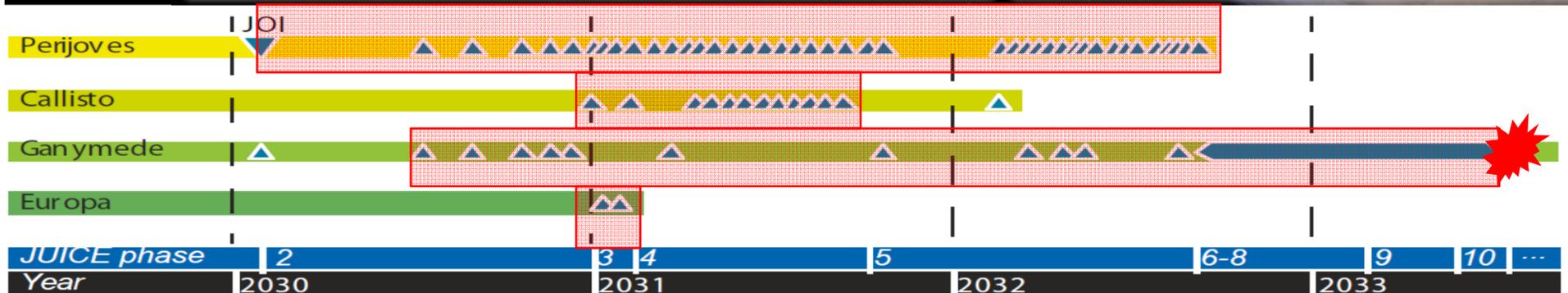
Ganymede:

planetary object and potential habitat

- Sub-surface, ice shell, ocean, interiors
- Geology, surface composition
- Atmosphere, ionosphere
- Magnetosphere, plasma environment

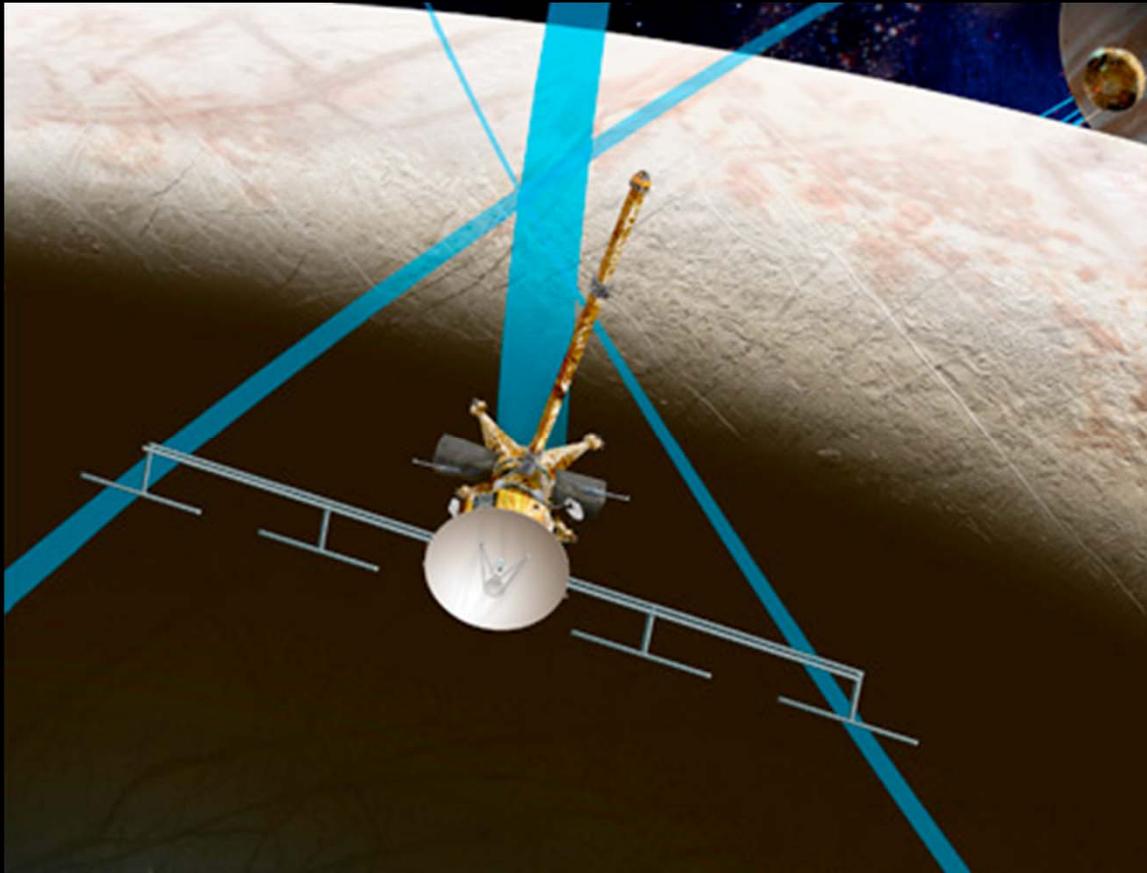
Jupiter System:

- Atmospheric structure, chemistry and dynamics
- Magnetosphere as fast rotator and giant accelerator
- Moons as plasma sources and sinks
- Couplings and interactions



Space missions

Outer system: the future projects cannot be dedicated to astrobiology



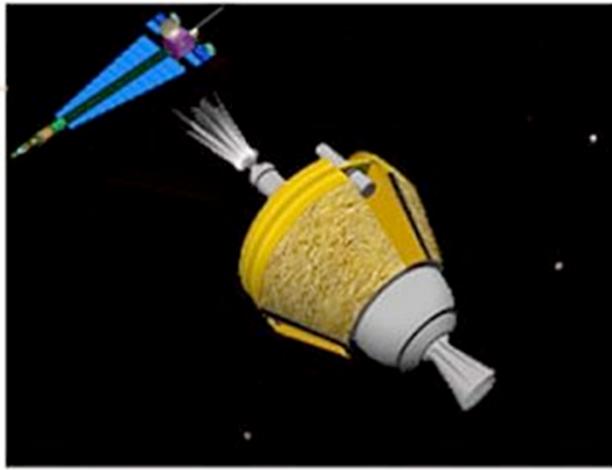
Europa Clipper



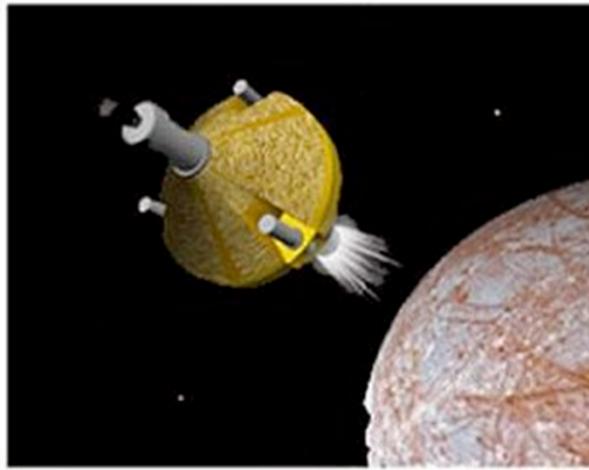
Europa lander: still a dream

Futures missions – quelques exemples vers le système externe

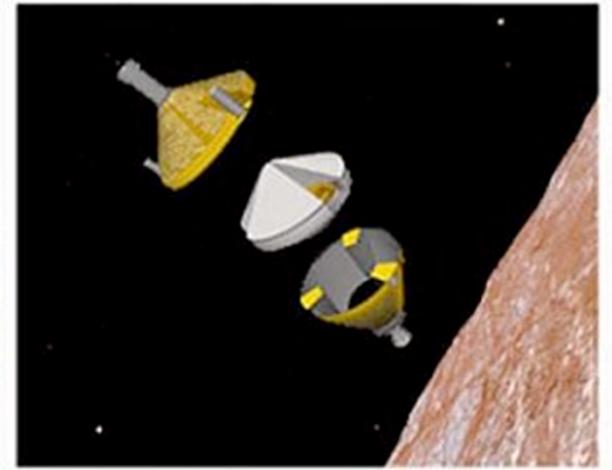
Europa Lander Mission (ELM) Separation, Entry and Landing Sequences



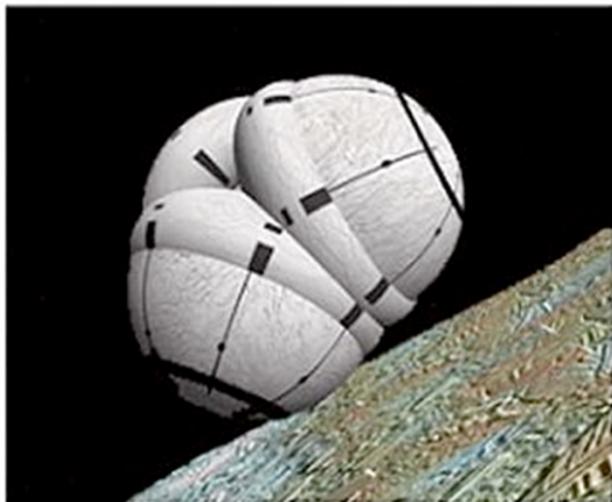
**Separation from JIMO
and Entry Burn #1 (Star 5)**



Entry Burn #2 (Star 17)



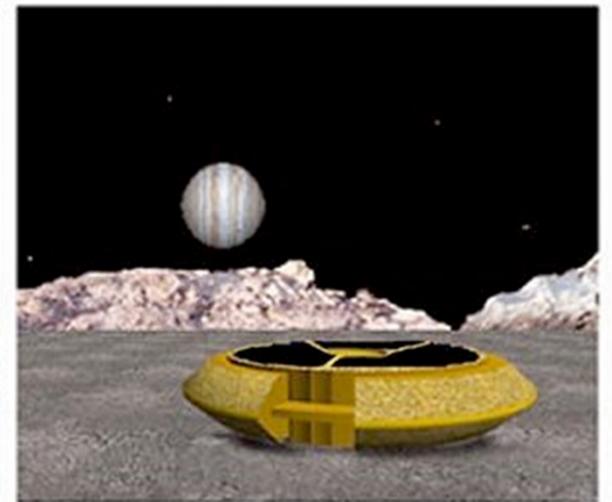
**Separation from
Propulsion Stages**



Descent



Deployment



Start 30 day Surface Mission

Space missions

Outer system: future projects to Titan and Enceladus

