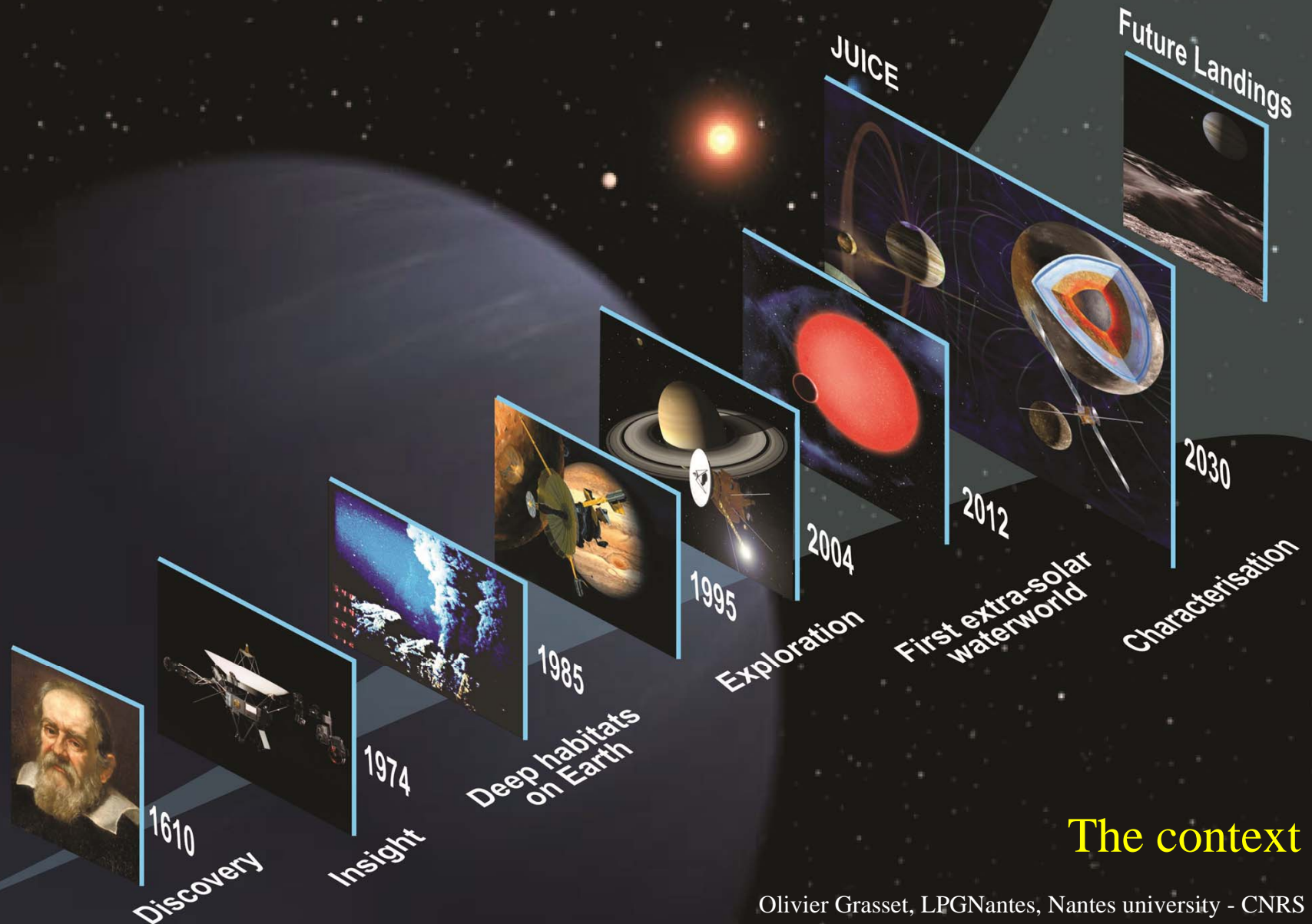


# Habitabilité des lunes de Jupiter: des premières évidences aux futures explorations

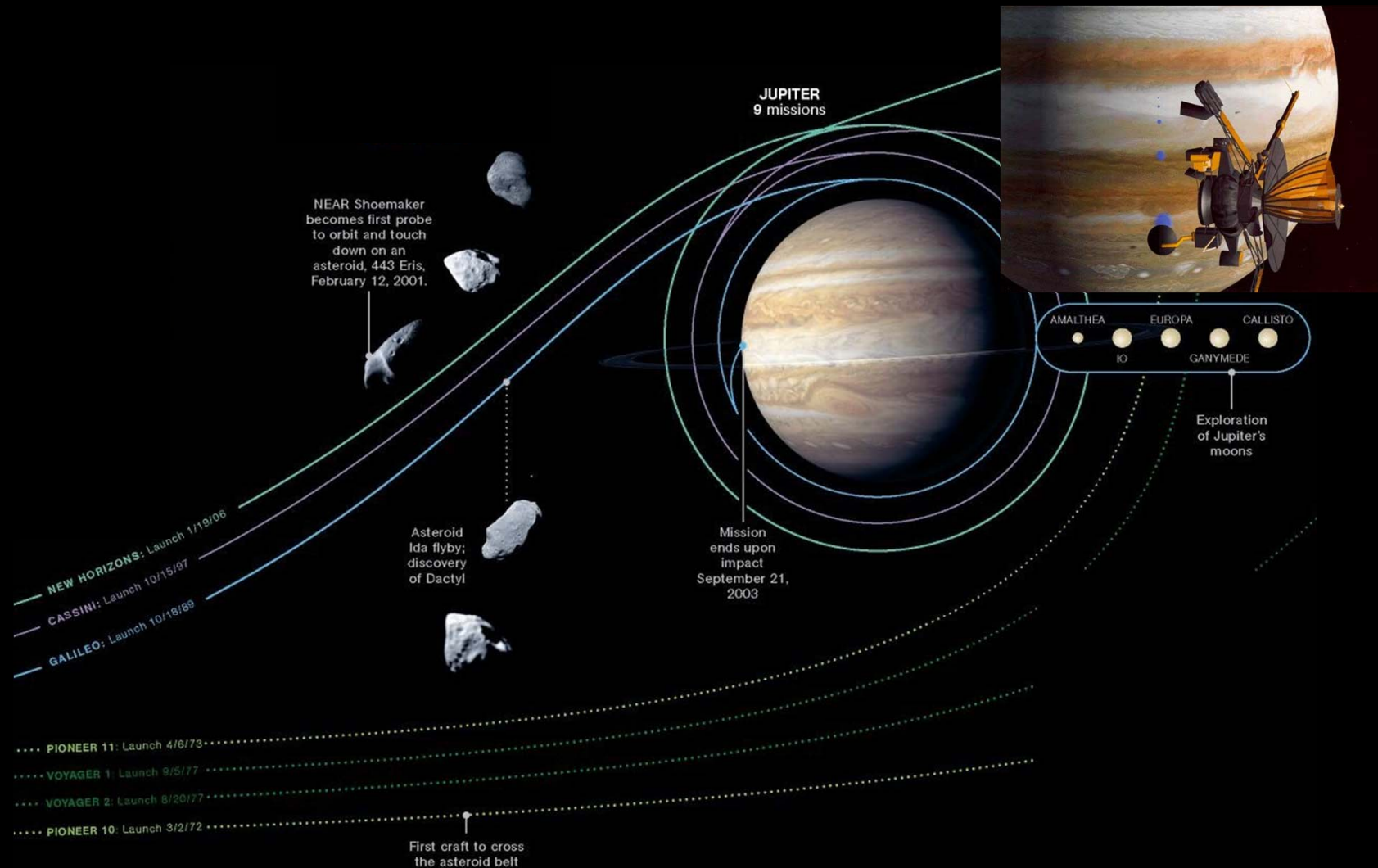


**The context**

# Previous missions

## Context

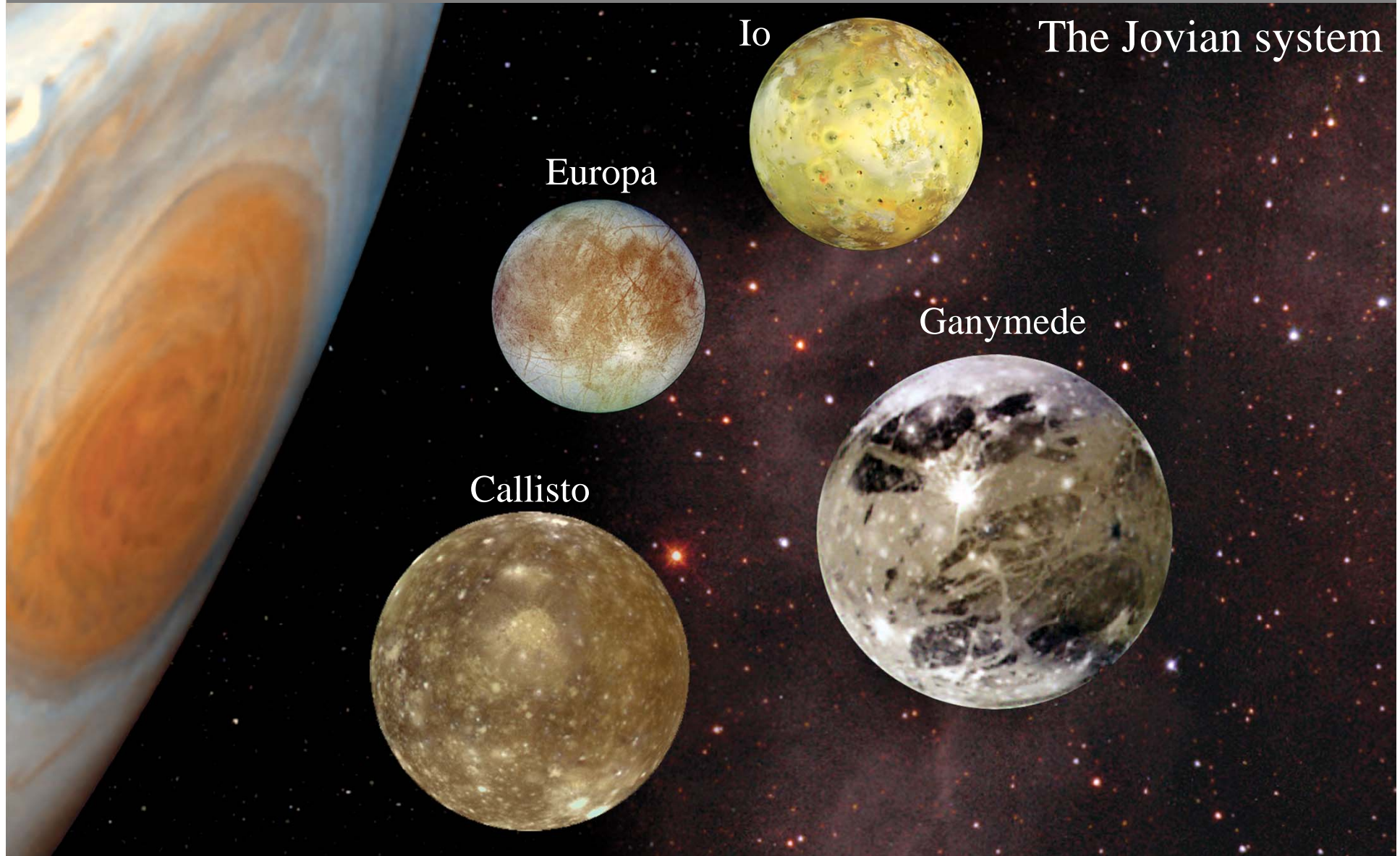
### A few flybys and an orbiter (Galileo)





# Habitabilité des lunes de Jupiter: des premières évidences aux futures explorations

The planetary bodies that we will explore...



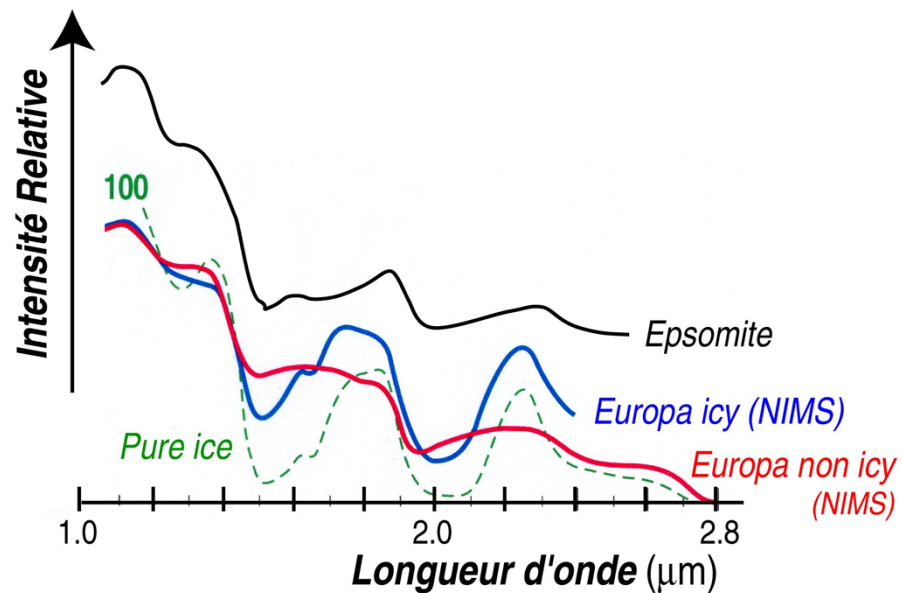


# The context

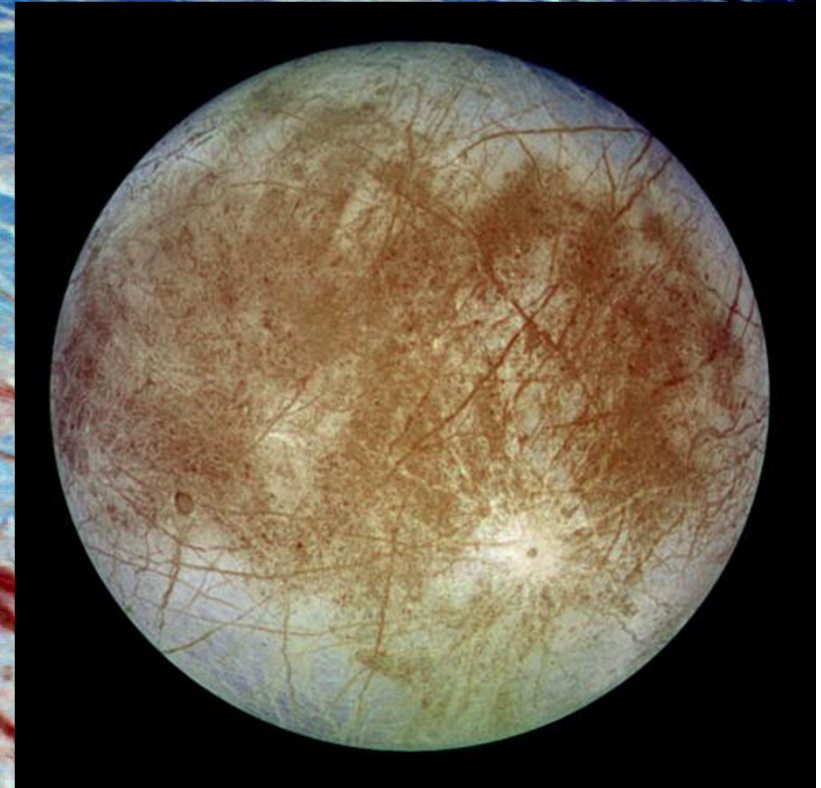
## Evidences of possible habitability

### Chemical evidences

### The composition of ices



from McCord et al. (1999)

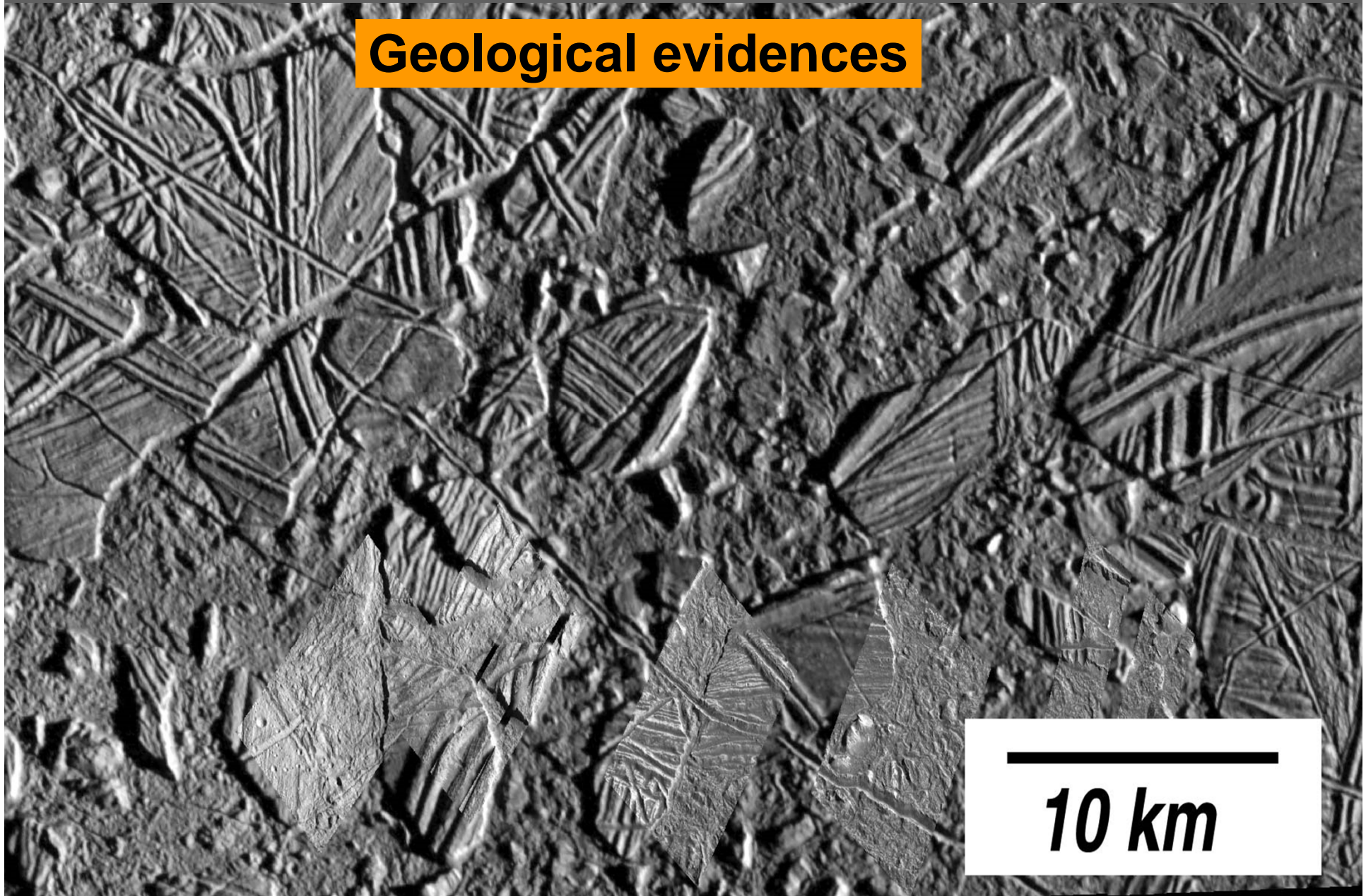




# The context

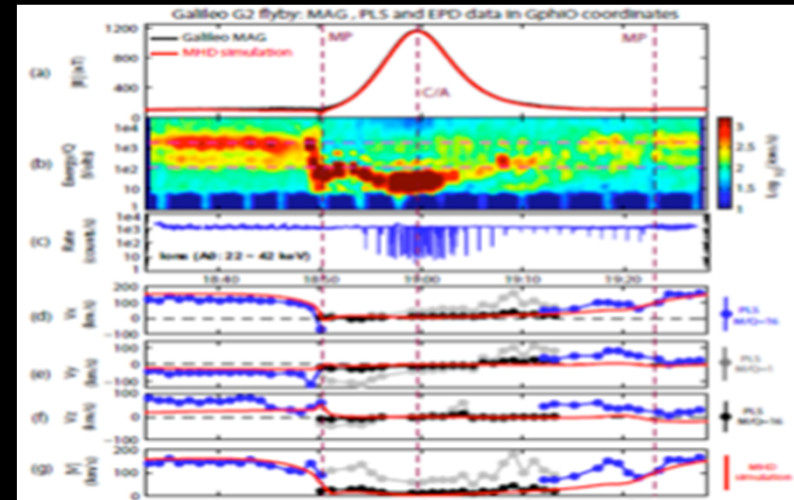
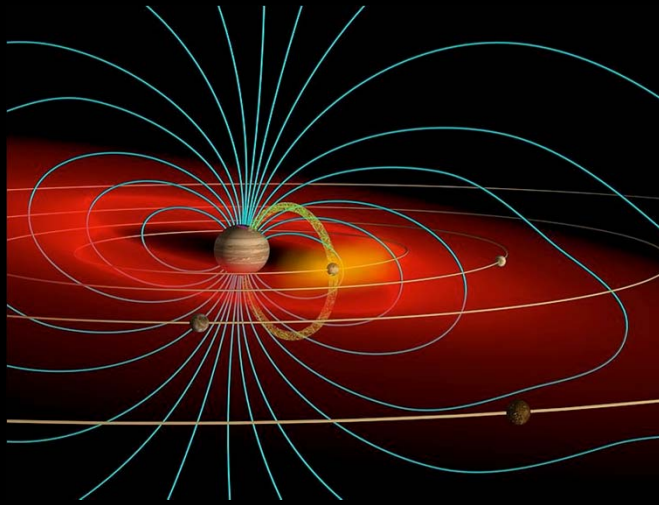
Evidences of possible habitability

## Geological evidences

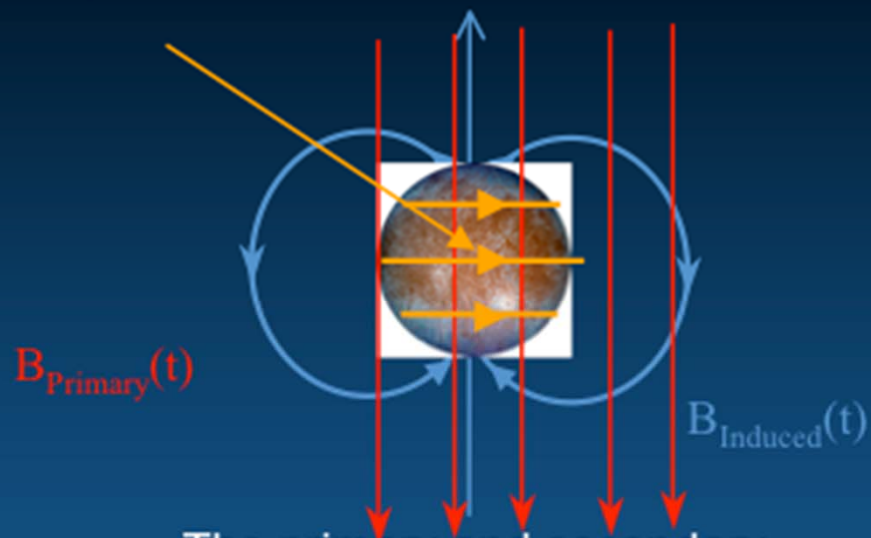


# The context

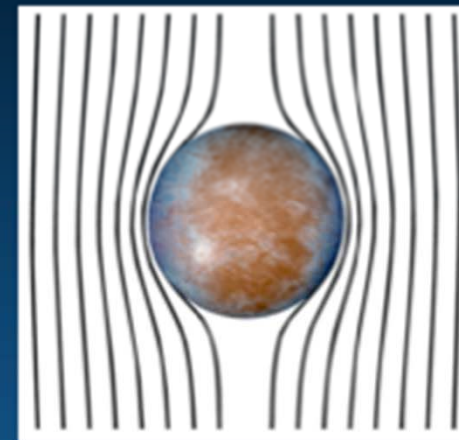
## Icy worlds – detection of an induced magnetic field



### Eddy currents



The primary and secondary fields shown separately

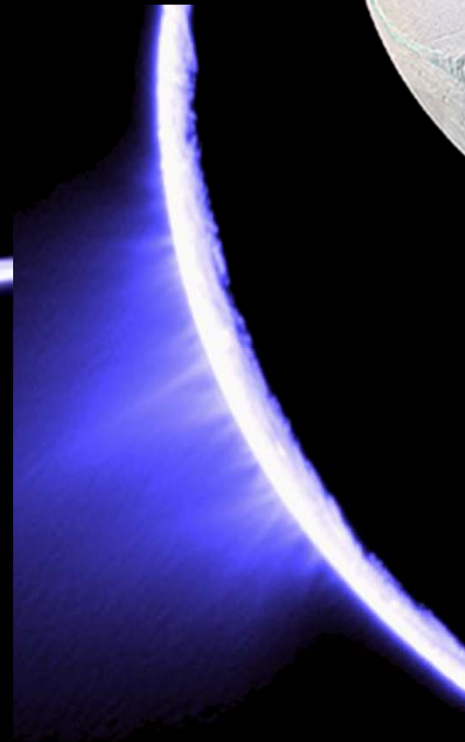
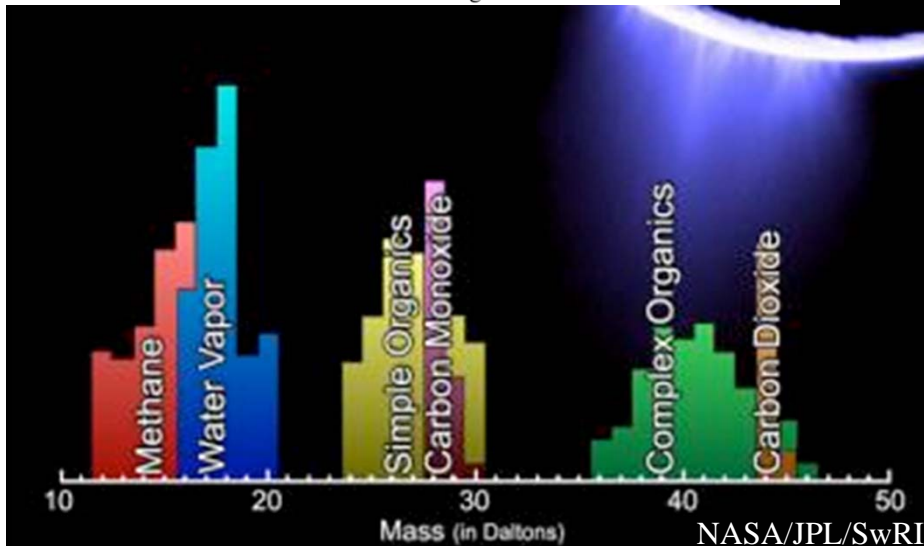
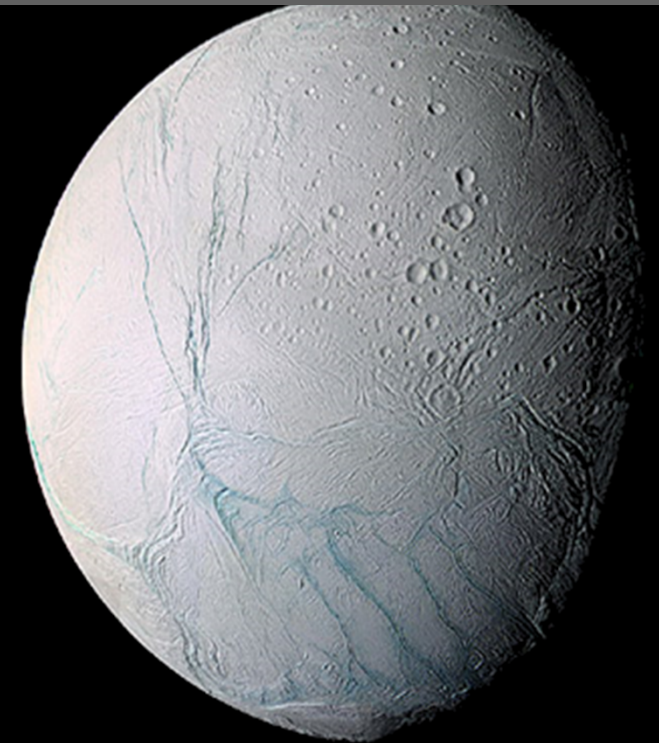
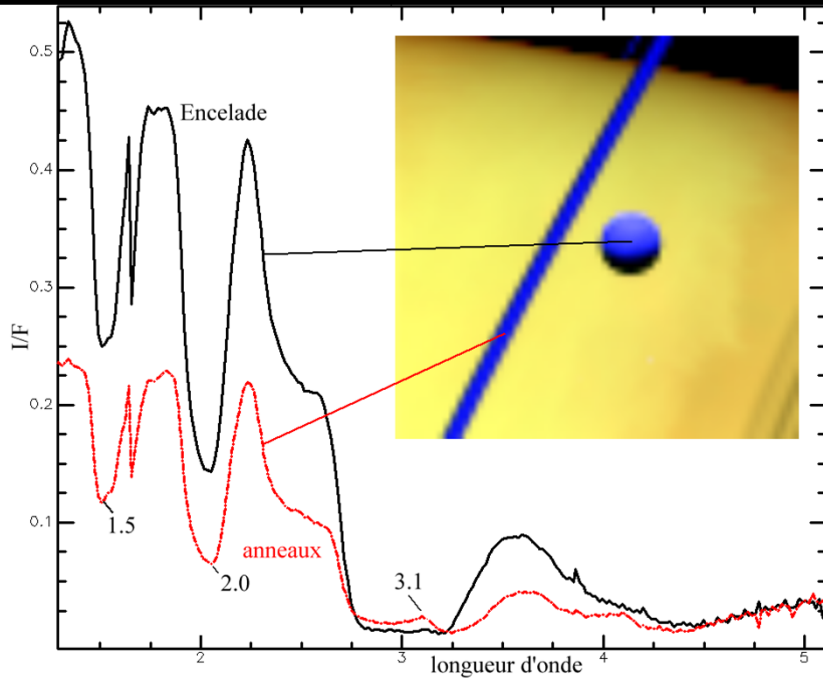


The total field



# The context

## Evidences of possible habitability – Cassini evidences at Enceladus



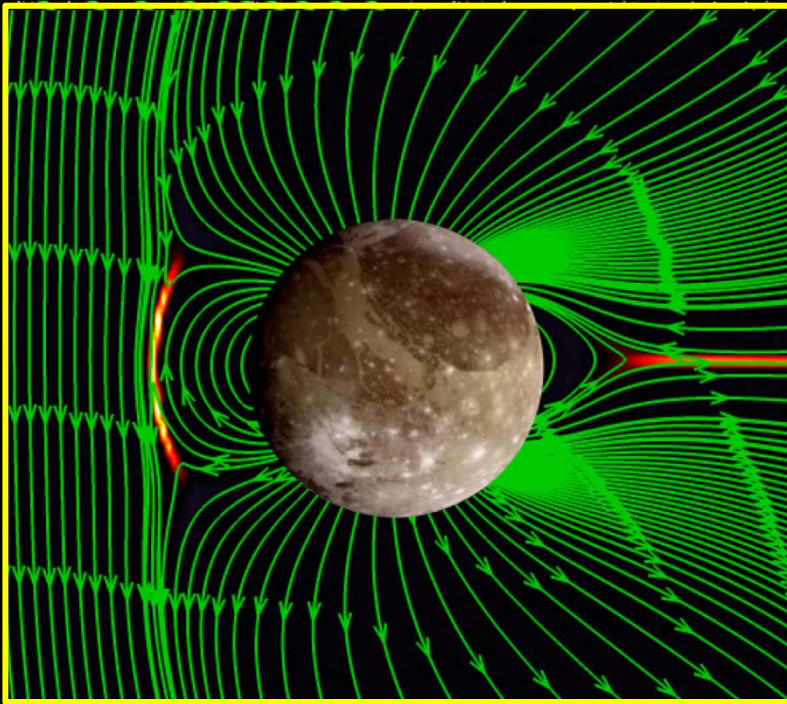
# The context

## Evidences of possible habitability – Ganymede and Callisto

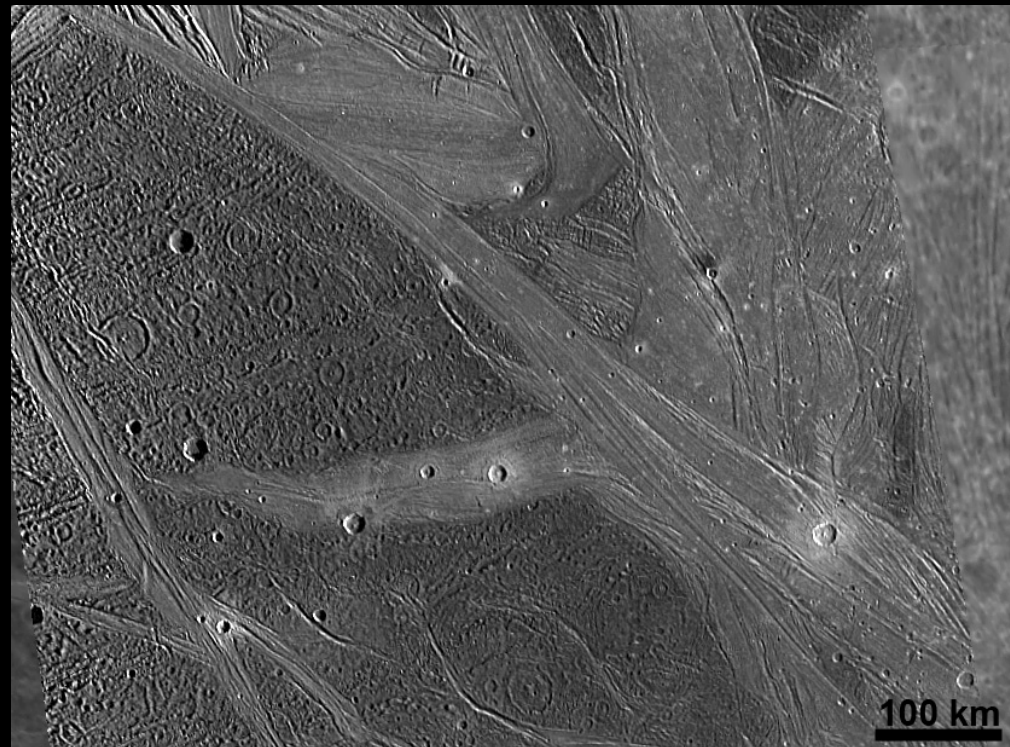
### Galileo evidences

#### Induced magnetic field

Observed but not characterised



#### Geologic activity

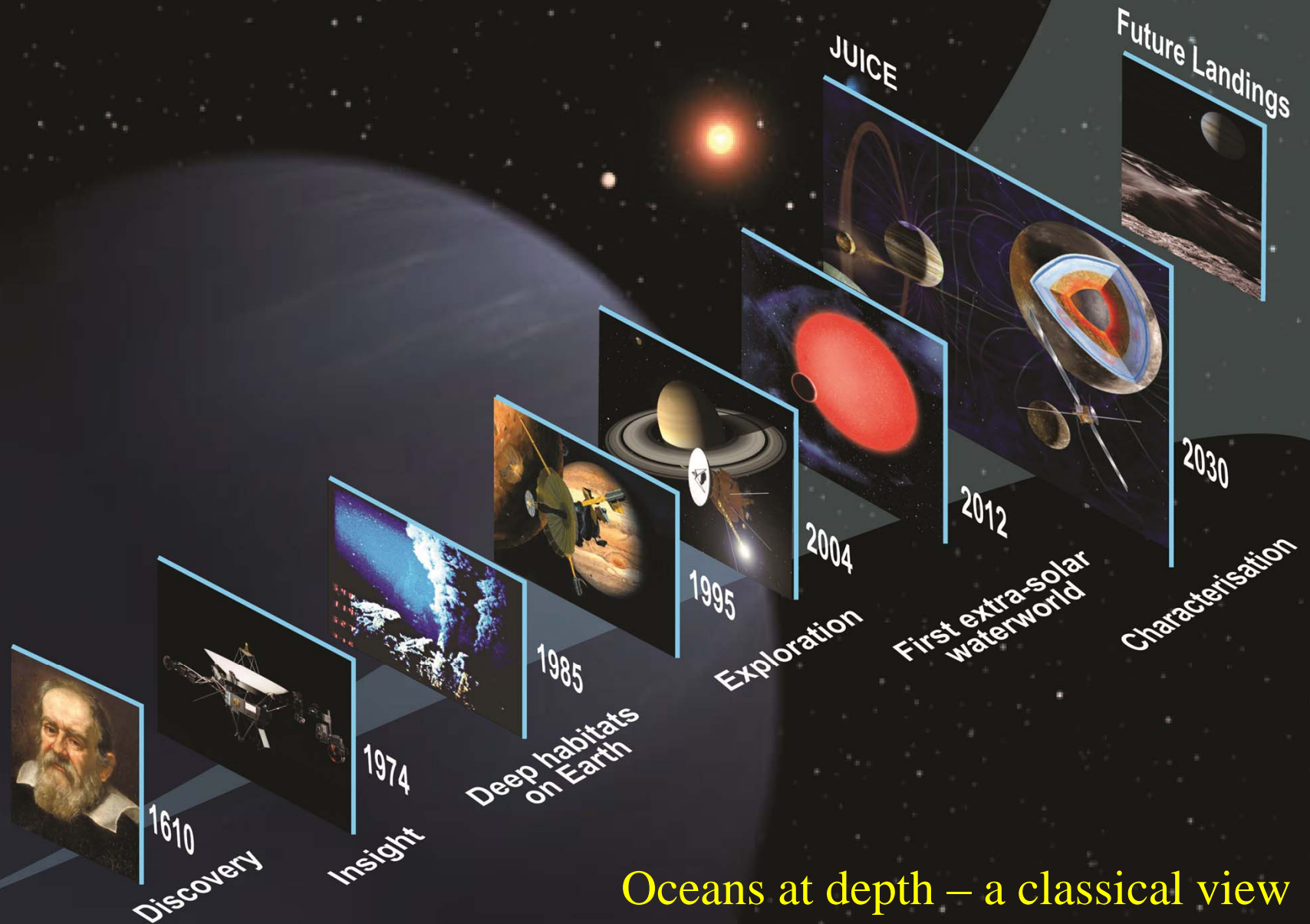


#### Questions

- ✧ Which depth?
- ✧ Which size?
- ✧ What is its composition?

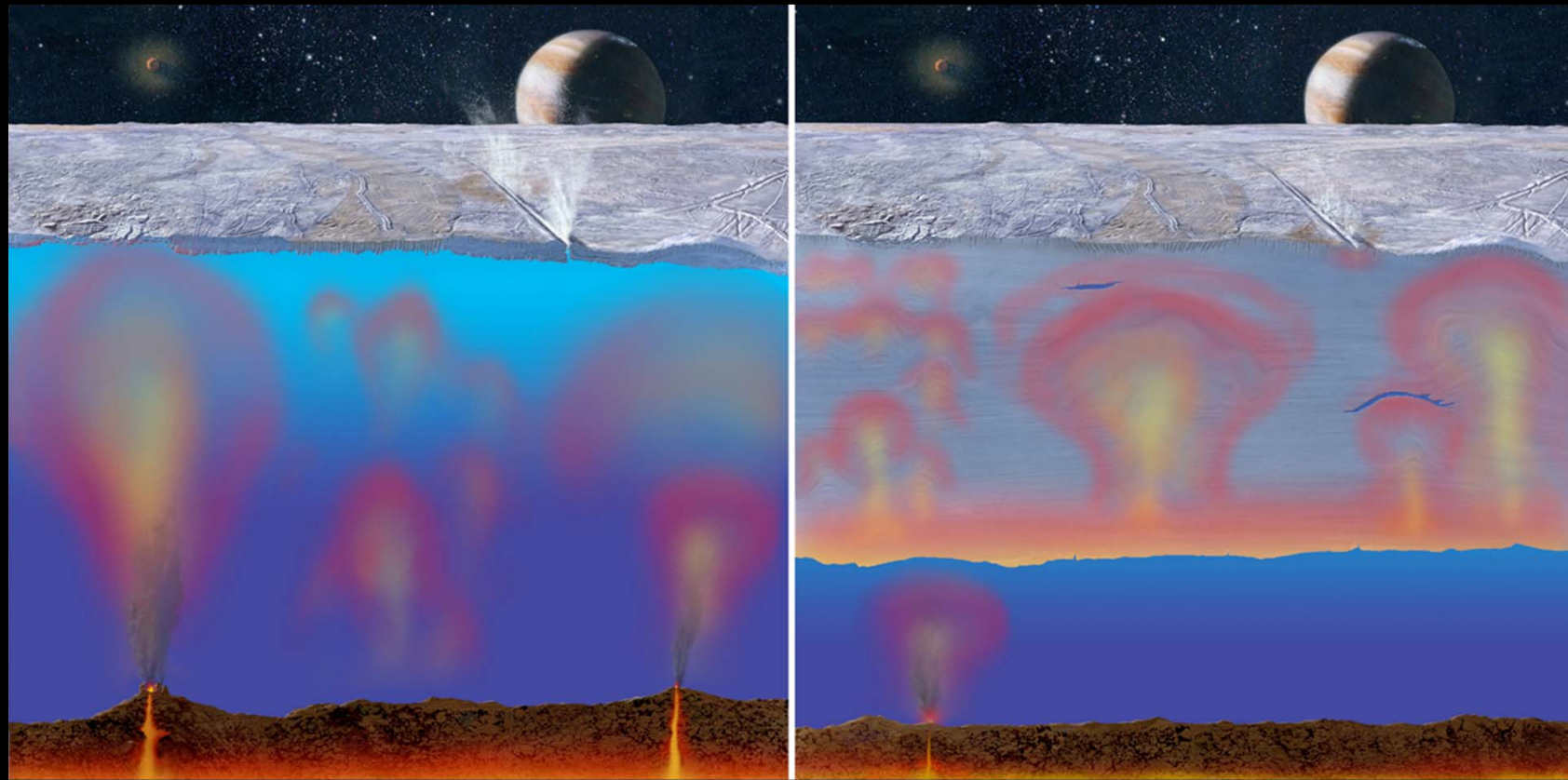


# Habitabilité des lunes de Jupiter: des premières évidences aux futures explorations



# Oceans at shallow depth

How deep are the oceans? We still don't know...



Europa

**Evidences are not sufficient to solve this issue**

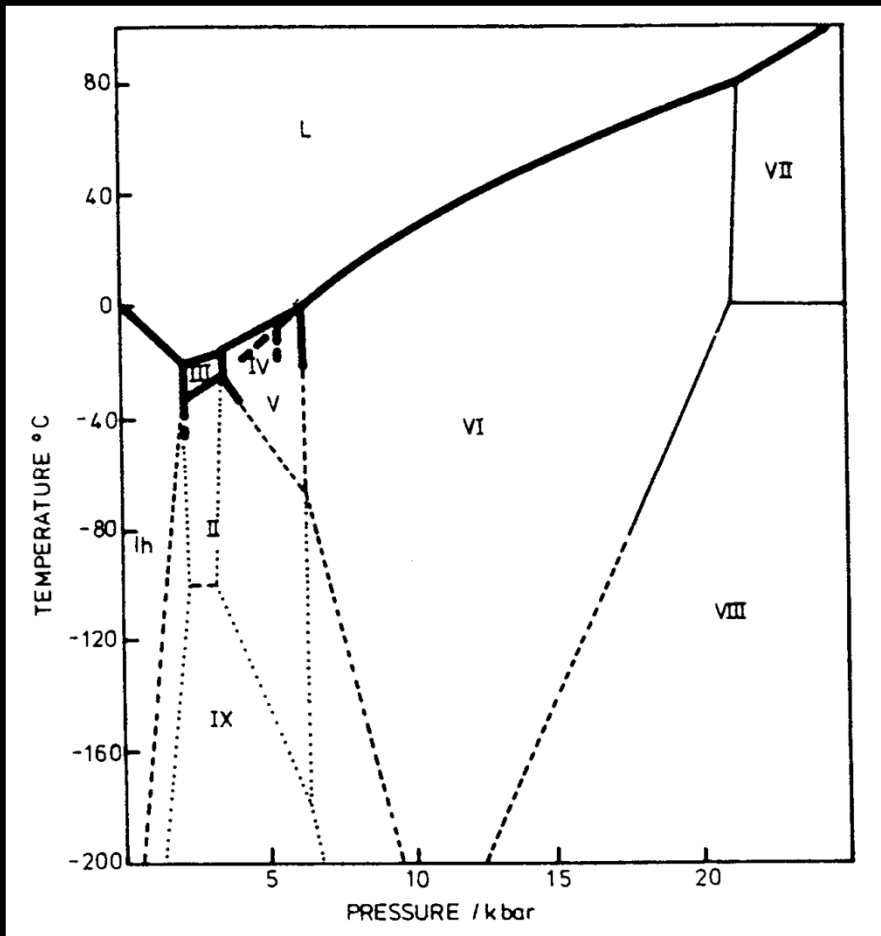
Globally thin crust: pros – magnetic field?, geologic features, current activity  
cons – thermal equilibrium?, geologic features

**A new space mission is needed**



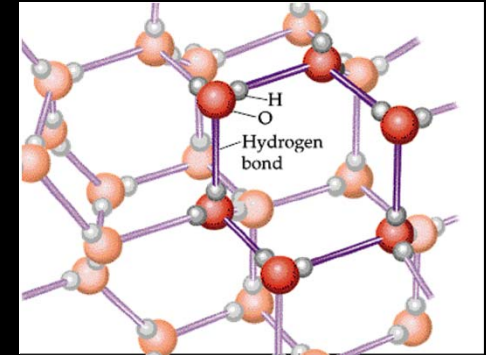
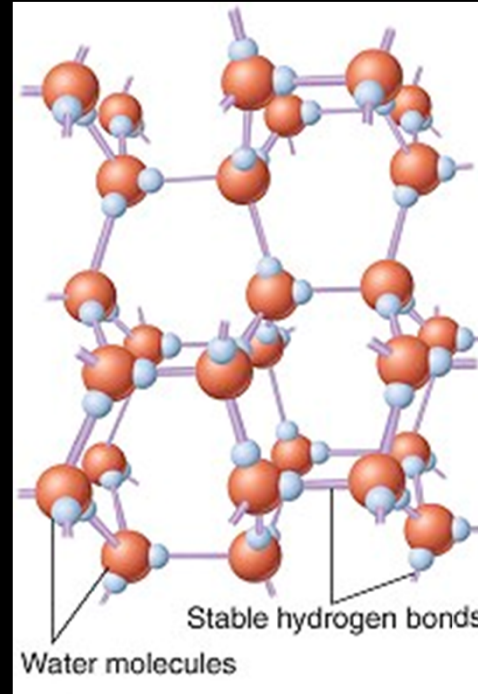
# Very deep oceans

## Giant Icy worlds – the high pressure phase diagram of water



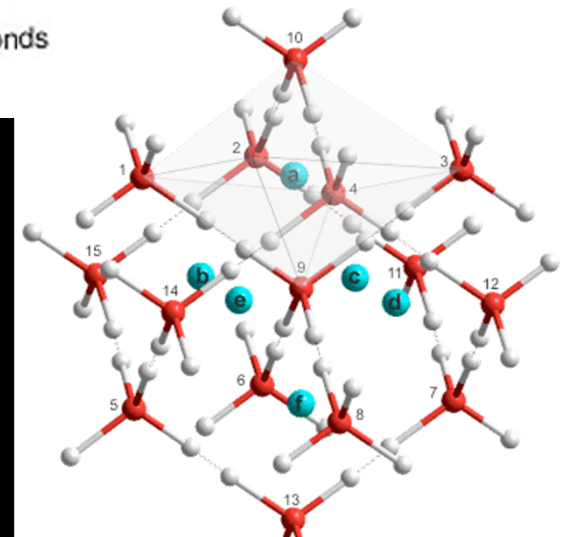
H<sub>2</sub>O Well-known since 1912 (Bridgman)  
Modern experiments (for planetology) devoted to complex mixtures.

Ice V



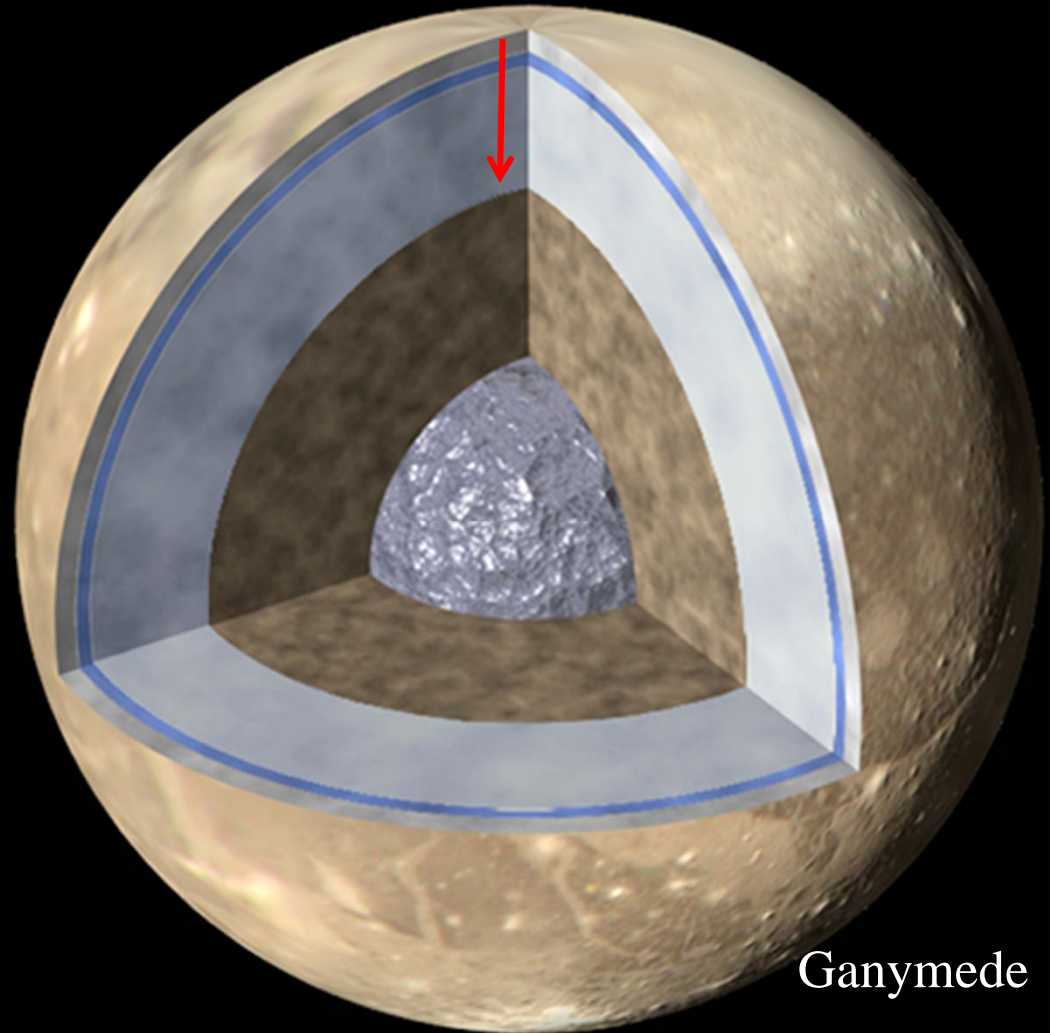
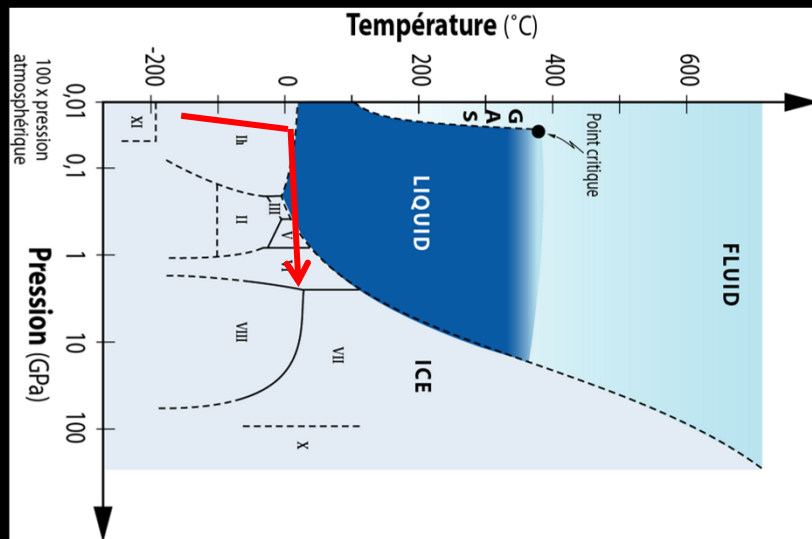
Ice Ih

Ice VII



# Very deep oceans

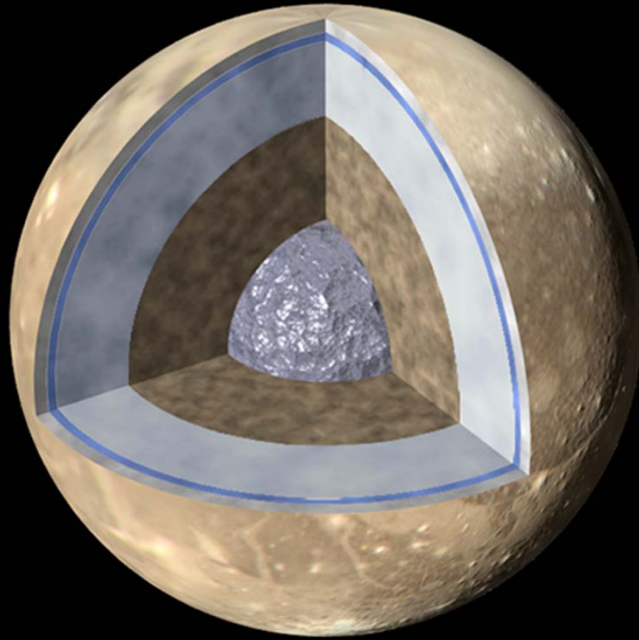
Icy worlds – evidences of a liquid ocean



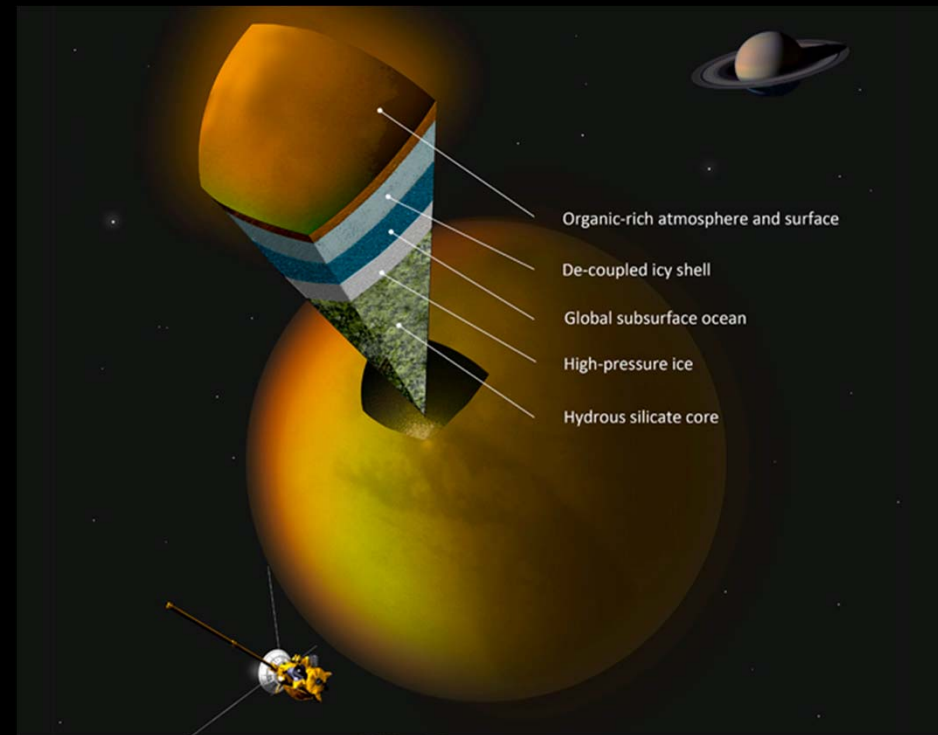


# Very deep oceans

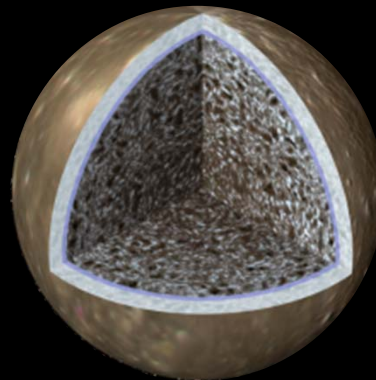
The known examples in the Solar System



Ganymede

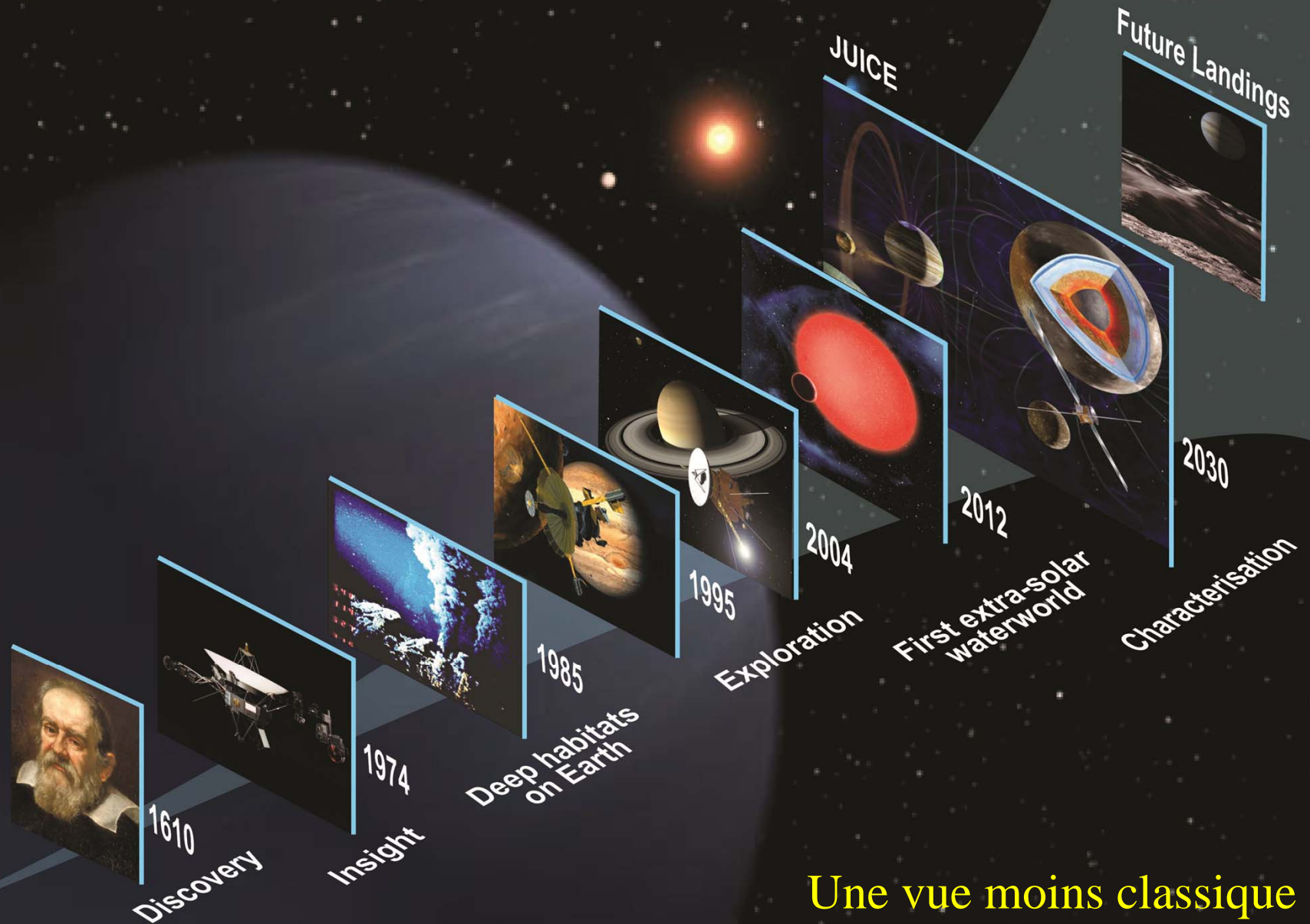


Titan



Callisto

# Habitabilité des lunes de Jupiter: des premières évidences aux futures explorations

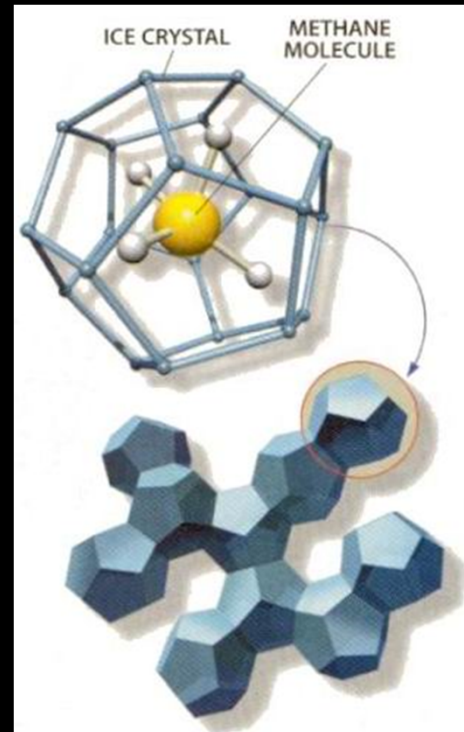
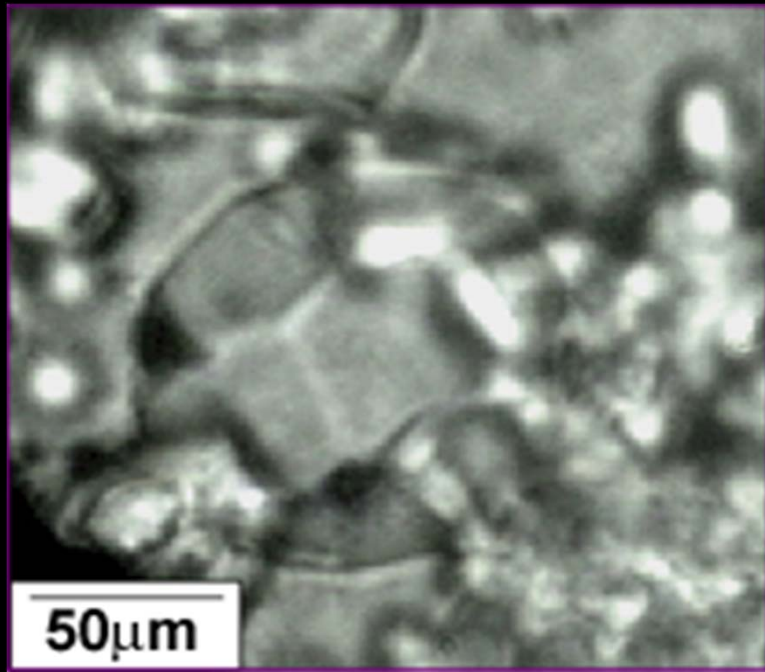




# Oceans at shallow depth

## Le problème de la composition

### *Clathrate hydrates : the burning ice.*



Known since the 50's (oil engineers)

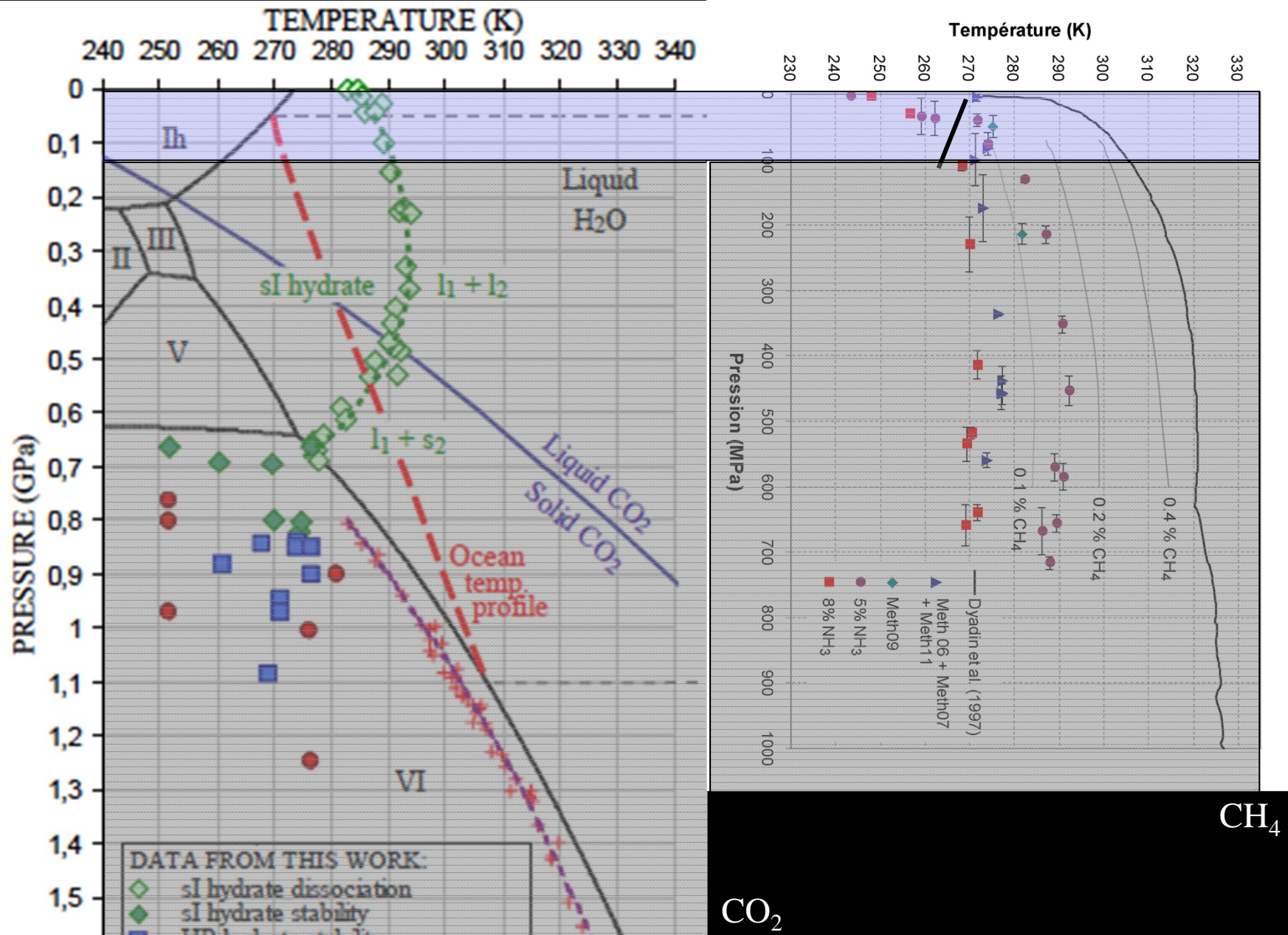
Envisaged in icy moons since early 90s (not sure about this...)

More recently, Prieto-Ballesteros (2005), and Vance et al. envisage to find them in abundance in Europa's crust

But how does that work in depth?

# Oceans at shallow depth

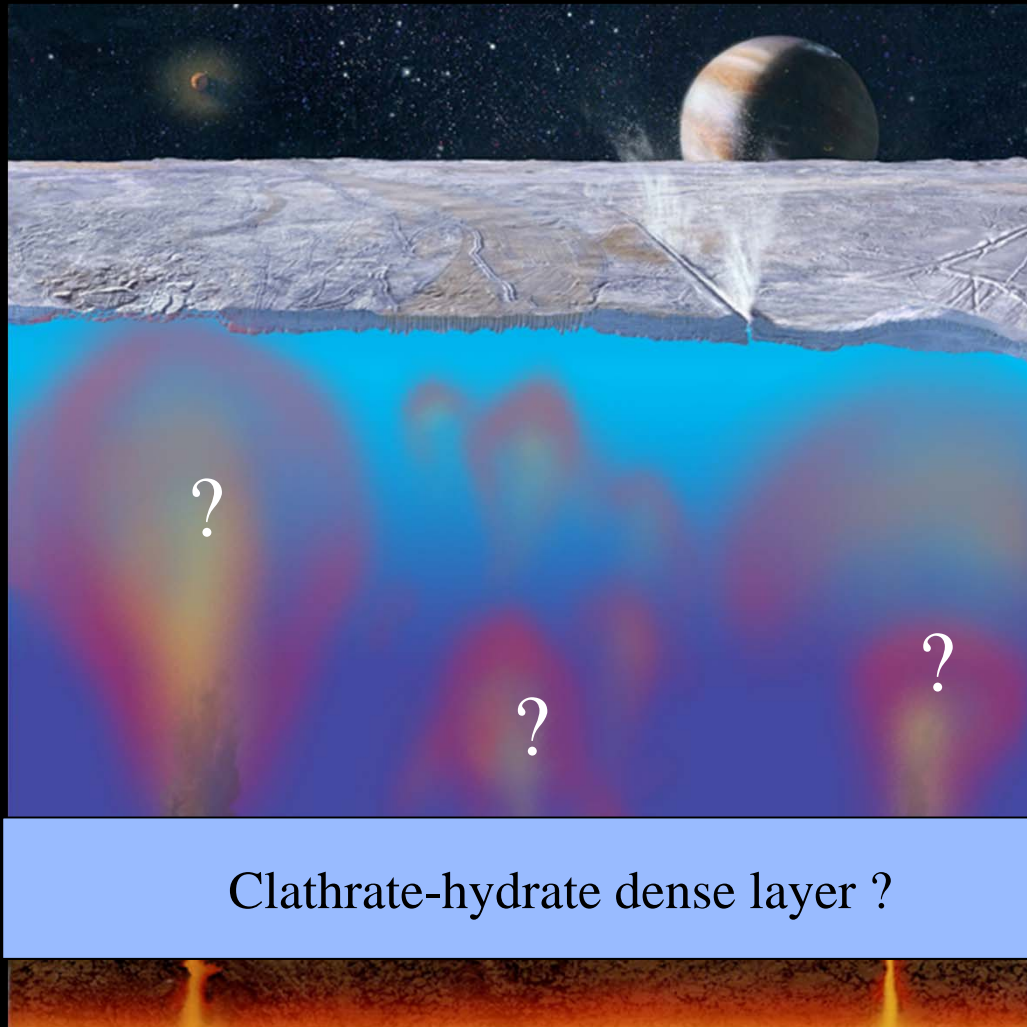
Contact of the ocean with silicates - is that certain?





# Oceans at shallow depth

Contact of the ocean with silicates - is that certain?



At last, the total amount of volatiles and salts fixes the thickness of the deep layer.

But why is that so important? A problem for **SPONCH**.

# Oceans at shallow depths

no Phosphorus in planetary ices

## Composition moléculaire

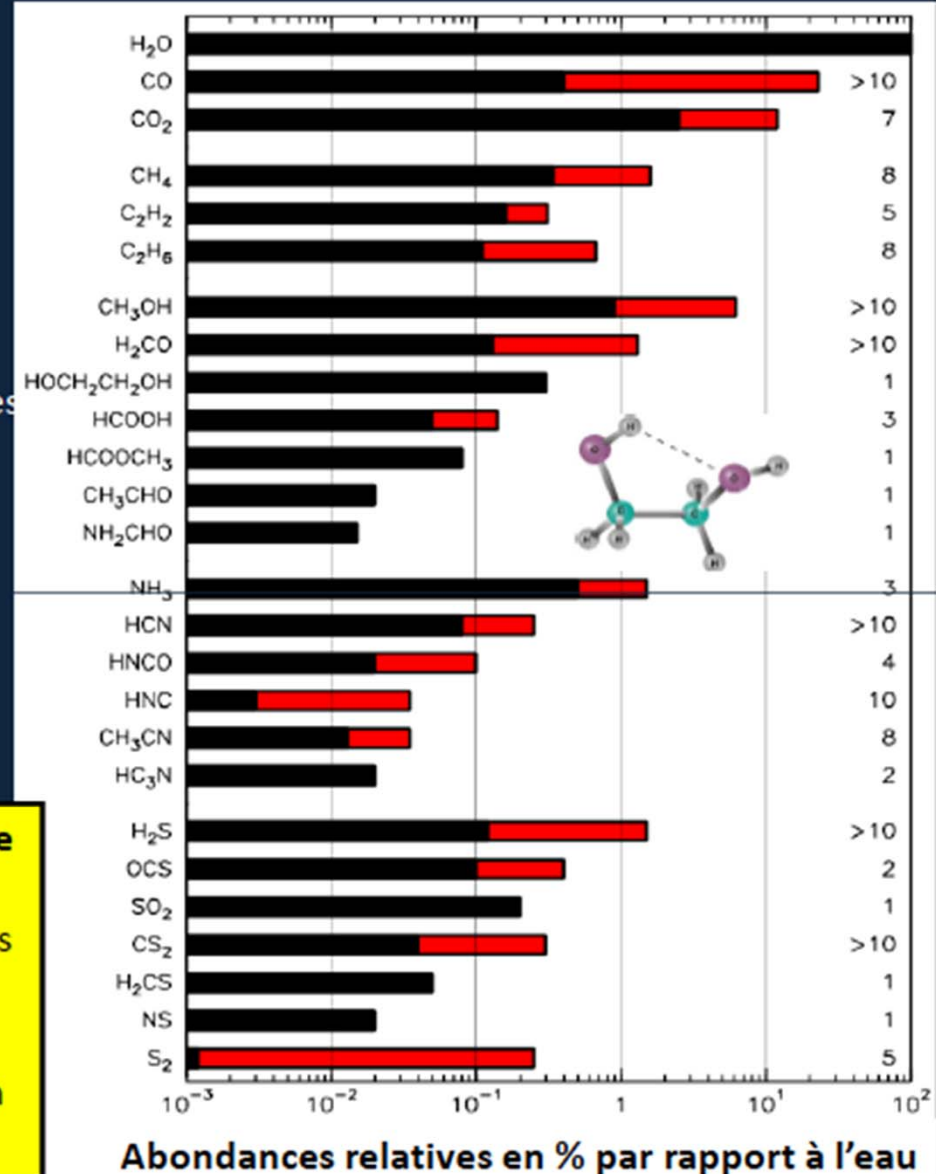
- grande richesse
- molécules complexes (ex: glycol)
- Molécules hydrogénées/oxygénées
- Molécules saturées/non saturées
- 5 ordres de grandeur dans les abondances

**Fortes similitudes avec les régions de formation d'étoiles**

**Formation par les mêmes mécanismes que les molécules interstellaires :**

- réactions ion-molécule, et à la surface des grains à basse température

**Origine :** grains interstellaires ou formation dans les régions extérieures du disque

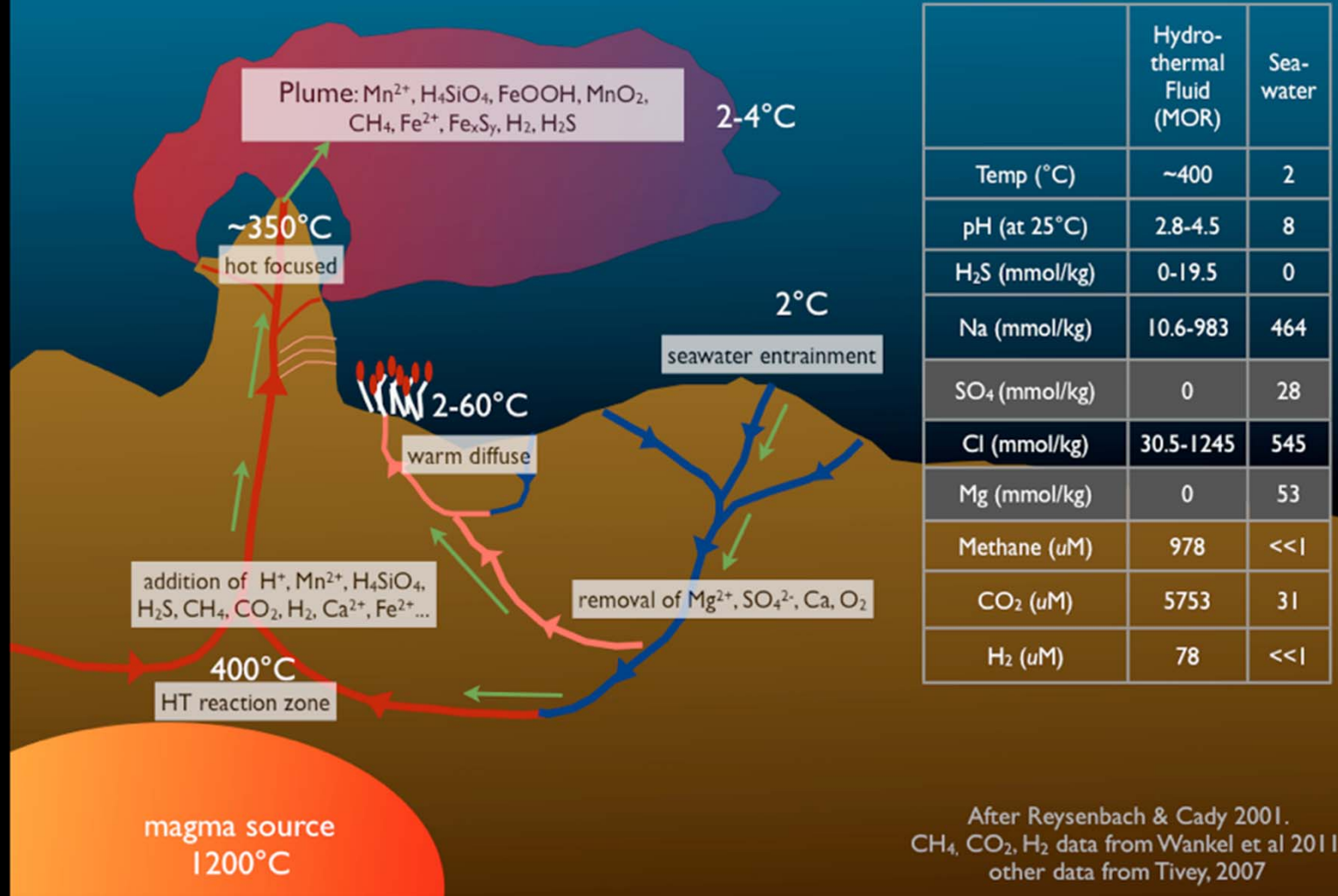




# Oceans at shallow depths

Hydrothermal vents – that helps a lot...

## Seawater changes into hydrothermal fluid

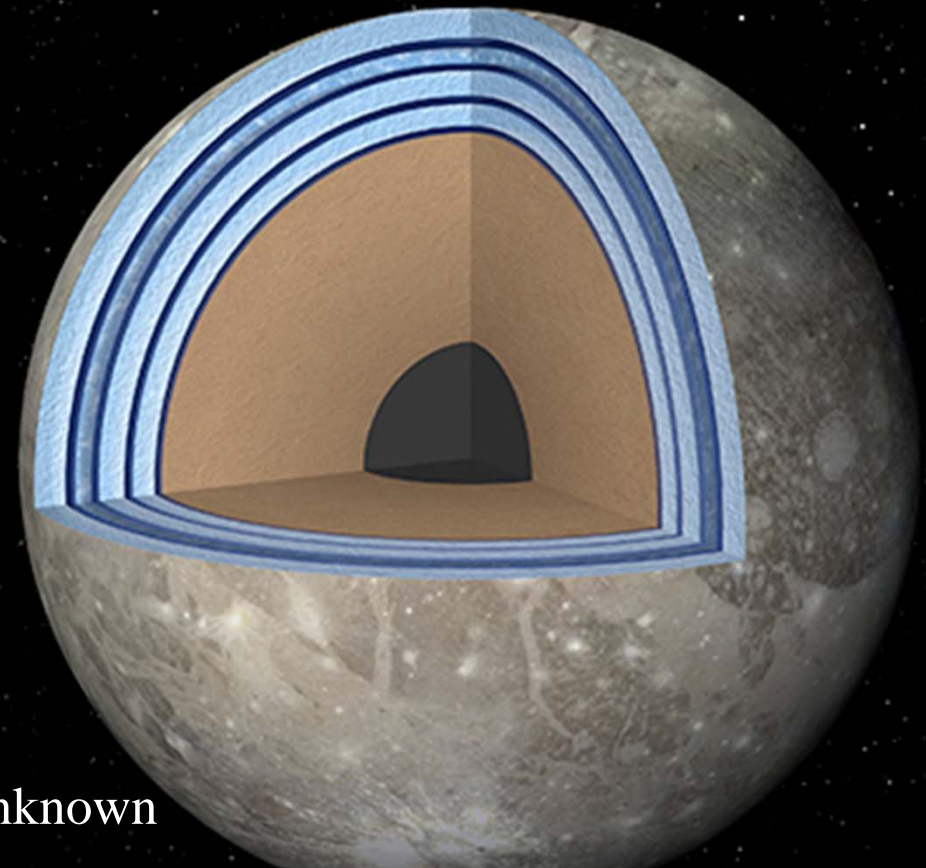
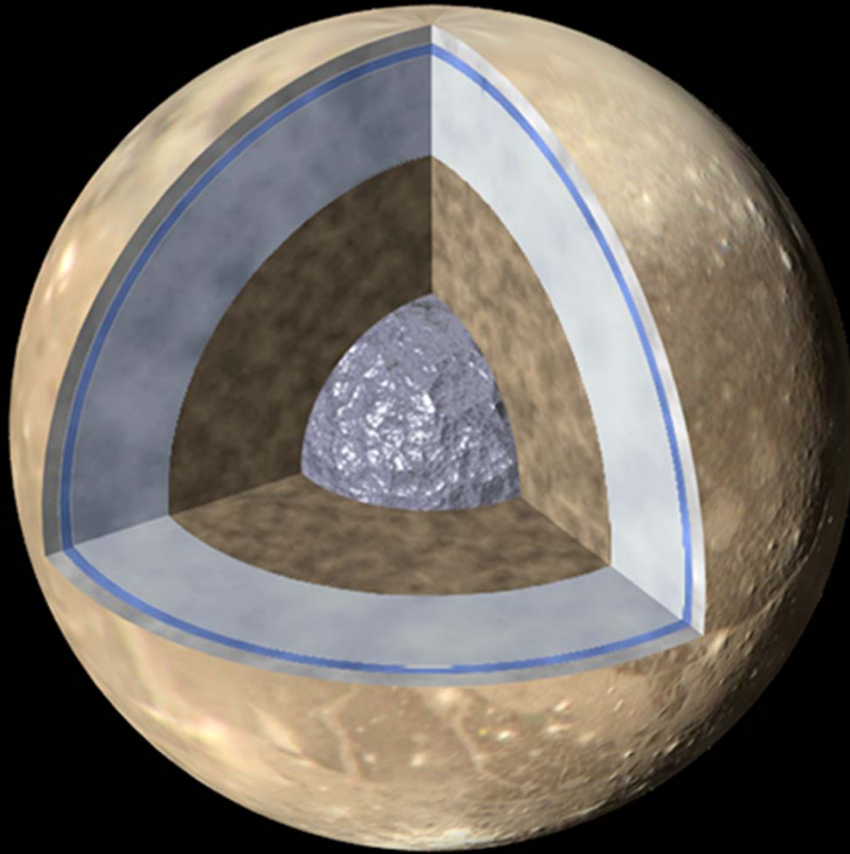


# Very deep oceans

Icy worlds – the complex chemistry is a key player ( $\text{NH}_3$ ,  $\text{H}_2\text{SO}_4$ ,  $\text{CO}_2$ ,  $\text{H}_2\text{S}$ , ...)

Density contrasts of water mixtures...

Vance et al., 2014



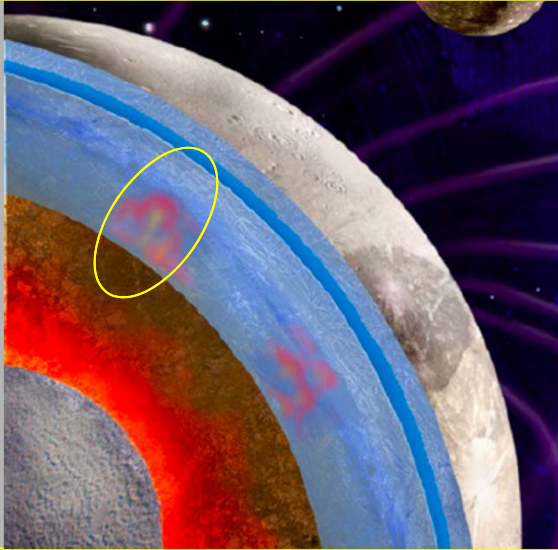
The internal structure of the moons is still unknown



# Very deep oceans

## A comparison

Ganymede type: liquid layers trapped between two icy mantles



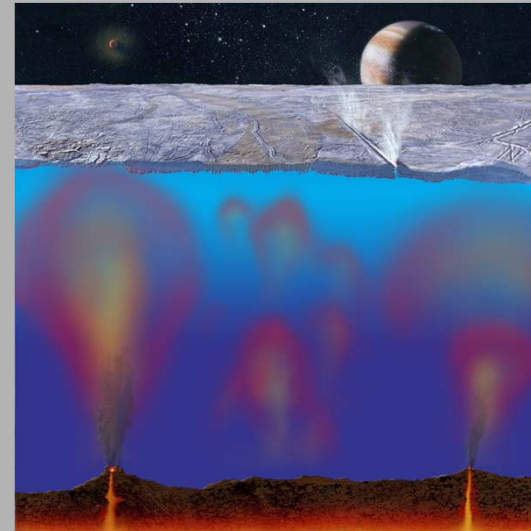
### Occurrence:

Largest moons, hot ice giants, ocean-planets...  
Most common habitat in the universe ?

### Key question:

Are these waterworlds habitable ?

Europa type: Liquid layer in direct contact with silicates (Earth's analog)



### Occurrence:

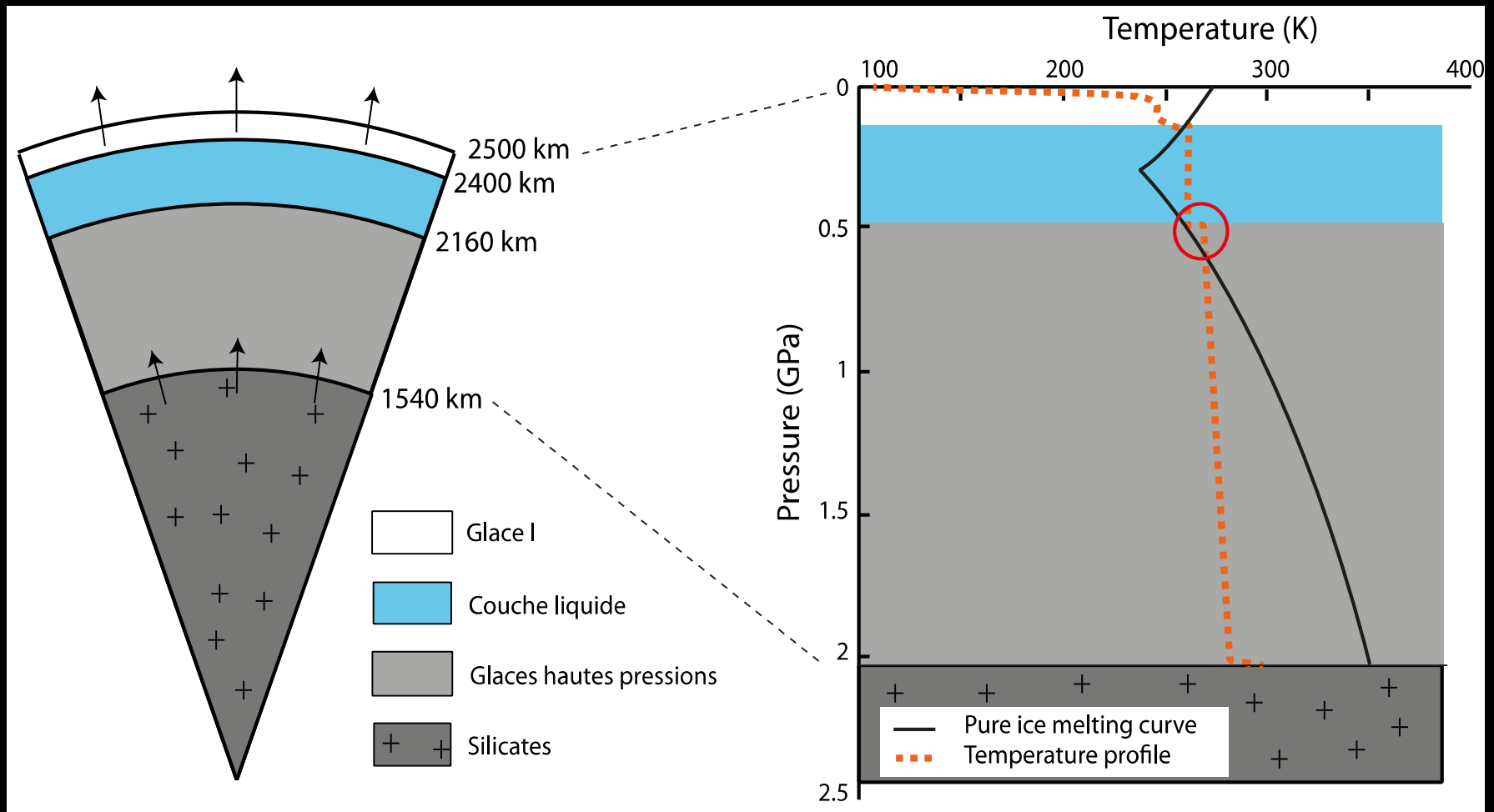
Europa, Enceladus  
Only possible for very small bodies

### Key questions:

How are the surface active areas related to potential deep habitats?  
Is the water in contact with silicates?

# Very deep oceans

## The dynamics of the system

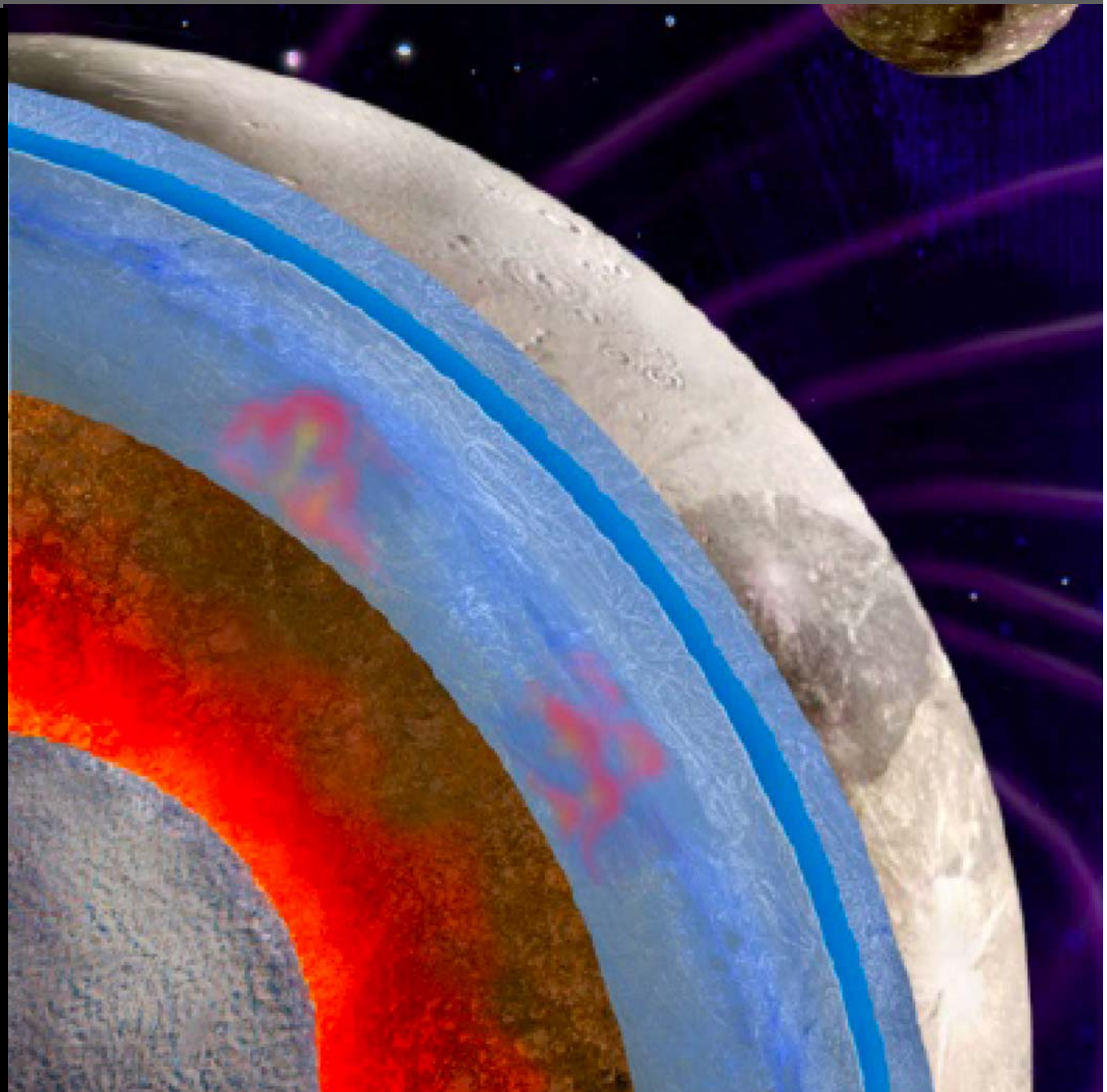


Convective regime

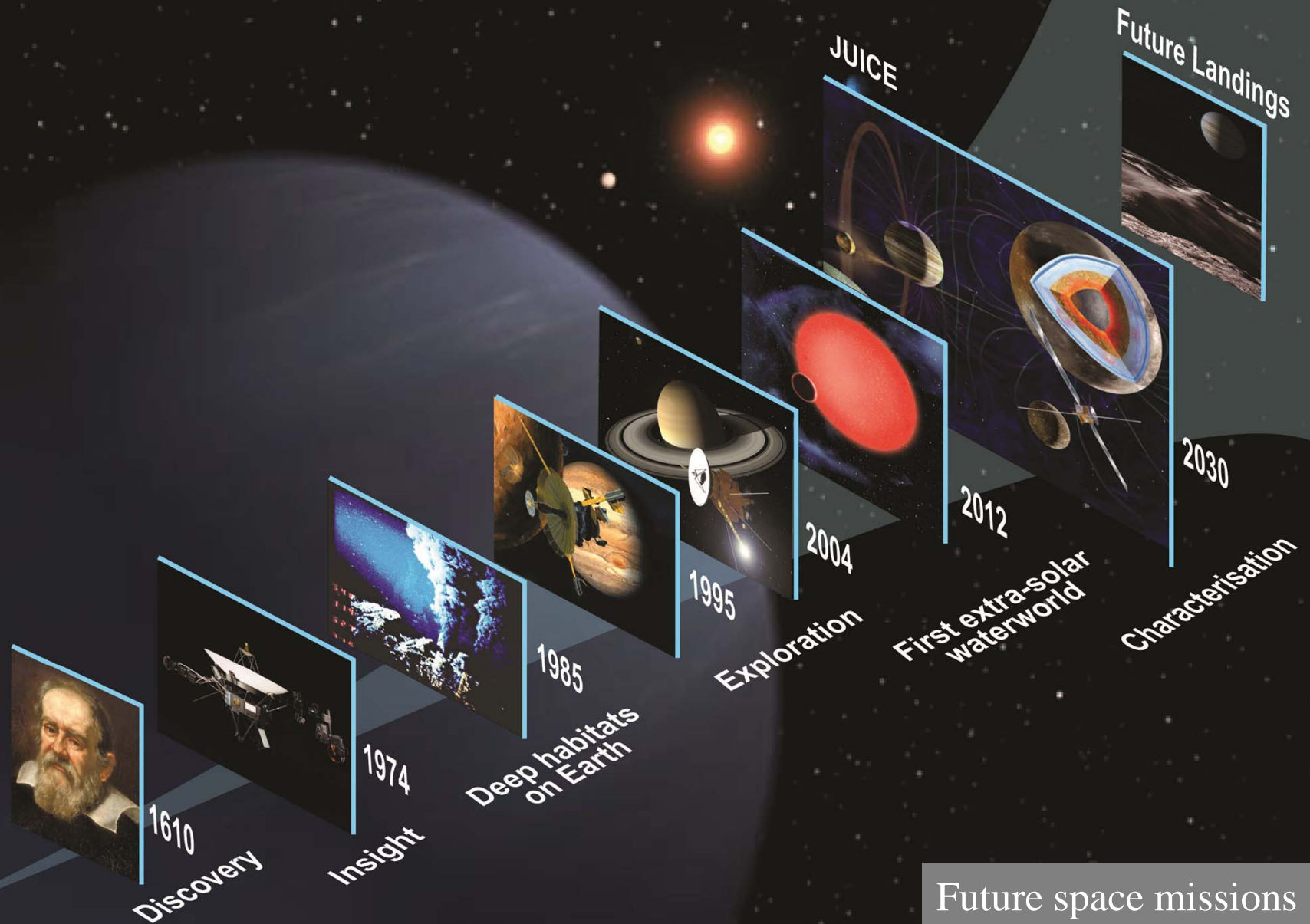
How does that work?  
Still unsolved



## Very deep oceans



# Habitabilité des lunes de Jupiter: des premières évidences aux futures explorations





# JUICE: JUpiter Icy moons Explorer

*Emergence of habitable worlds around gas giants*

*Jupiter system as an archetype for gas giants*

## Callisto:

*remnant of the early solar system*

- Icy shell, ocean
- Geology, surface composition
- Past activity

## Europa: recently active zones

- Surface non-water-ice material
- Search for liquid water
- Recent activity

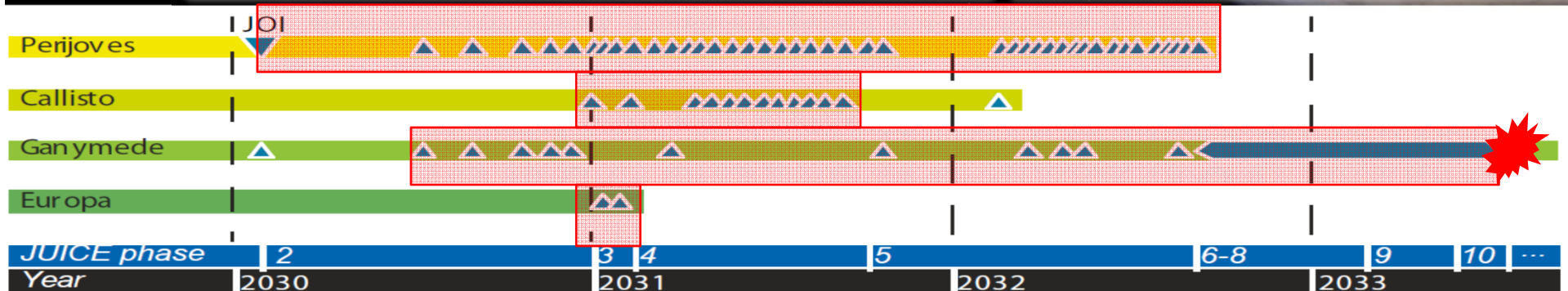
## Jupiter System:

- Atmospheric structure, chemistry and dynamics
- Magnetosphere as fast rotator and giant accelerator
- Moons as plasma sources and sinks
- Couplings and interactions

## Ganymede:

*planetary object and potential habitat*

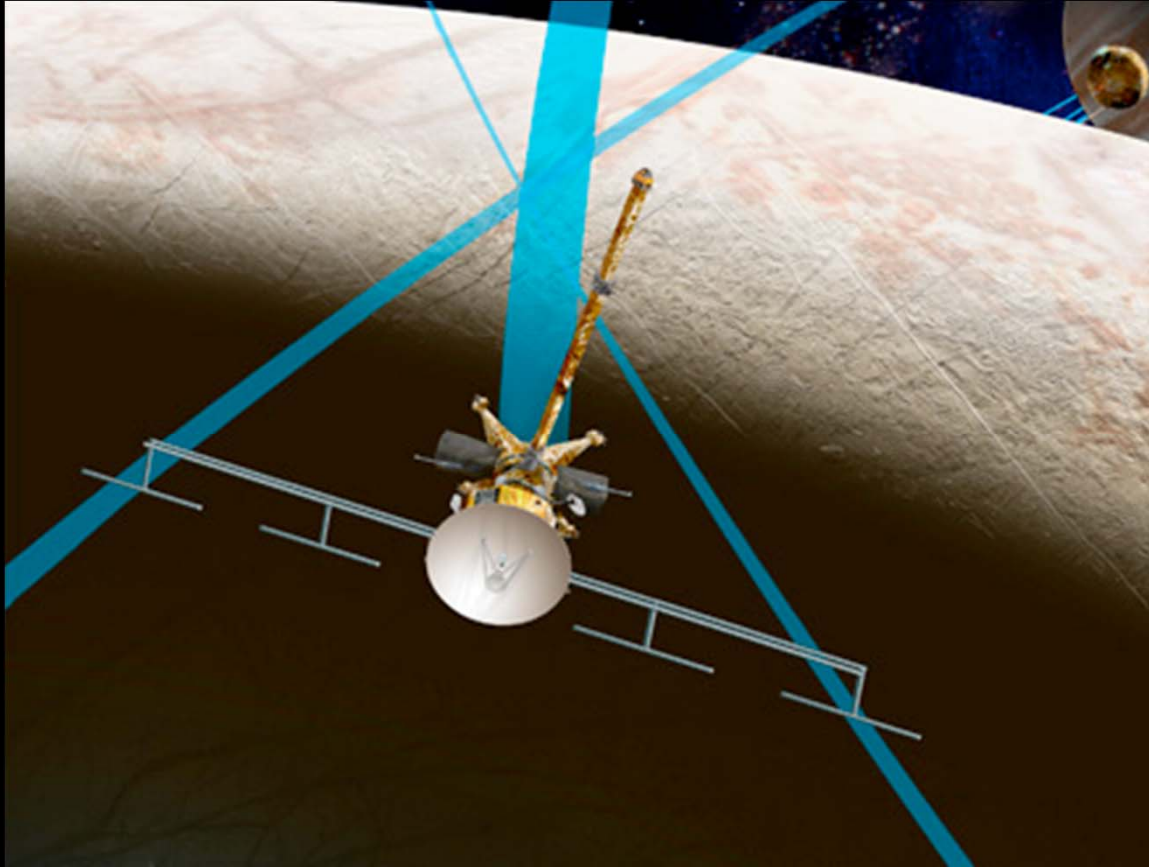
- Sub-surface, ice shell, ocean, interiors
- Geology, surface composition
- Atmosphere, ionosphere
- Magnetosphere, plasma environment





# Space missions

Outer system: the future projects cannot be dedicated to astrobiology



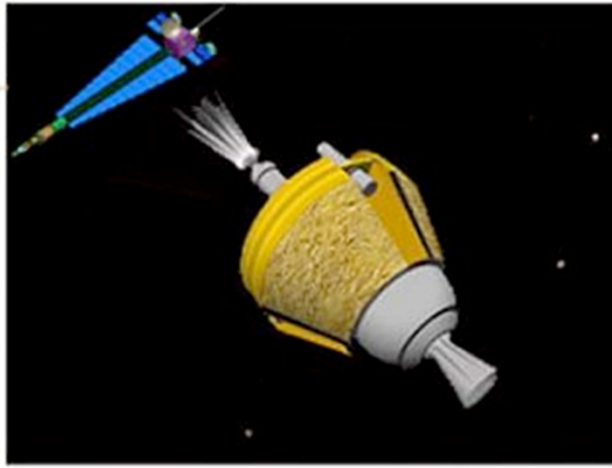
Europa Clipper



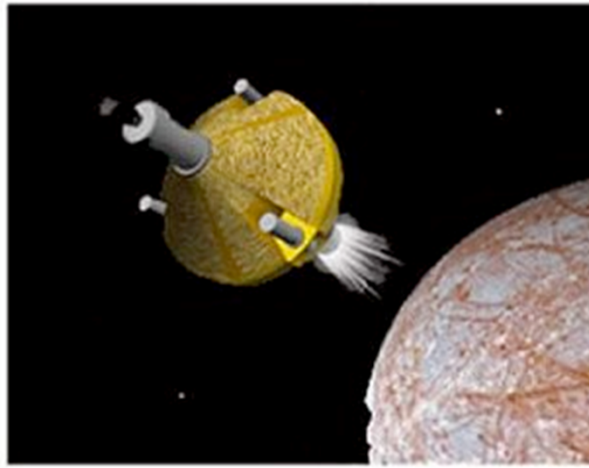
Europa lander: still a dream

## Futures missions – quelques exemples vers le système externe

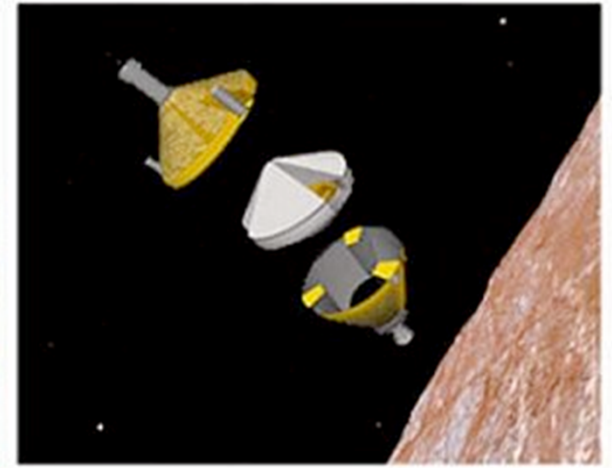
### Europa Lander Mission (ELM) Separation, Entry and Landing Sequences



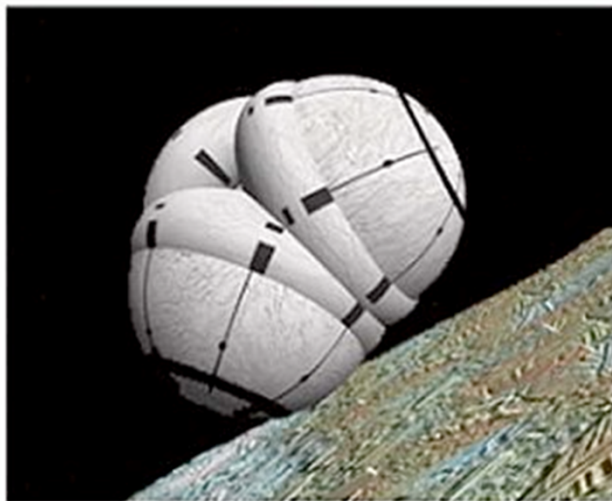
Separation from JIMO  
and Entry Burn #1 (Star 5)



Entry Burn #2 (Star 17)



Separation from  
Propulsion Stages



Descent



Deployment



Start 30 day Surface Mission

# Space missions

Outer system: future projects to Titan and Enceladus

